

Bayesian inference

Paul Lasky

**“Understanding the
underlying processes that
create data streams is a
ubiquitous problem across the
sciences”**

Matilda B. Bilby*

***not a real quote
(also not a real Bilby)**



- 1% of women at age forty have breast cancer.
- 80% of women with breast cancer have positive mammograms.
- 9.6% of women without breast cancer also get positive mammograms (false positive).
- A woman in this age group had a positive mammogram in a routine screening. What is the probability that she has breast cancer?

probability of picking a random 40-yr old woman that has cancer:

$$p(A) = \frac{|A|}{|W|}$$

probability of picking a random 40-yr old woman that will have a positive mammogram (with or without cancer):

$$p(B) = \frac{|B|}{|W|}$$

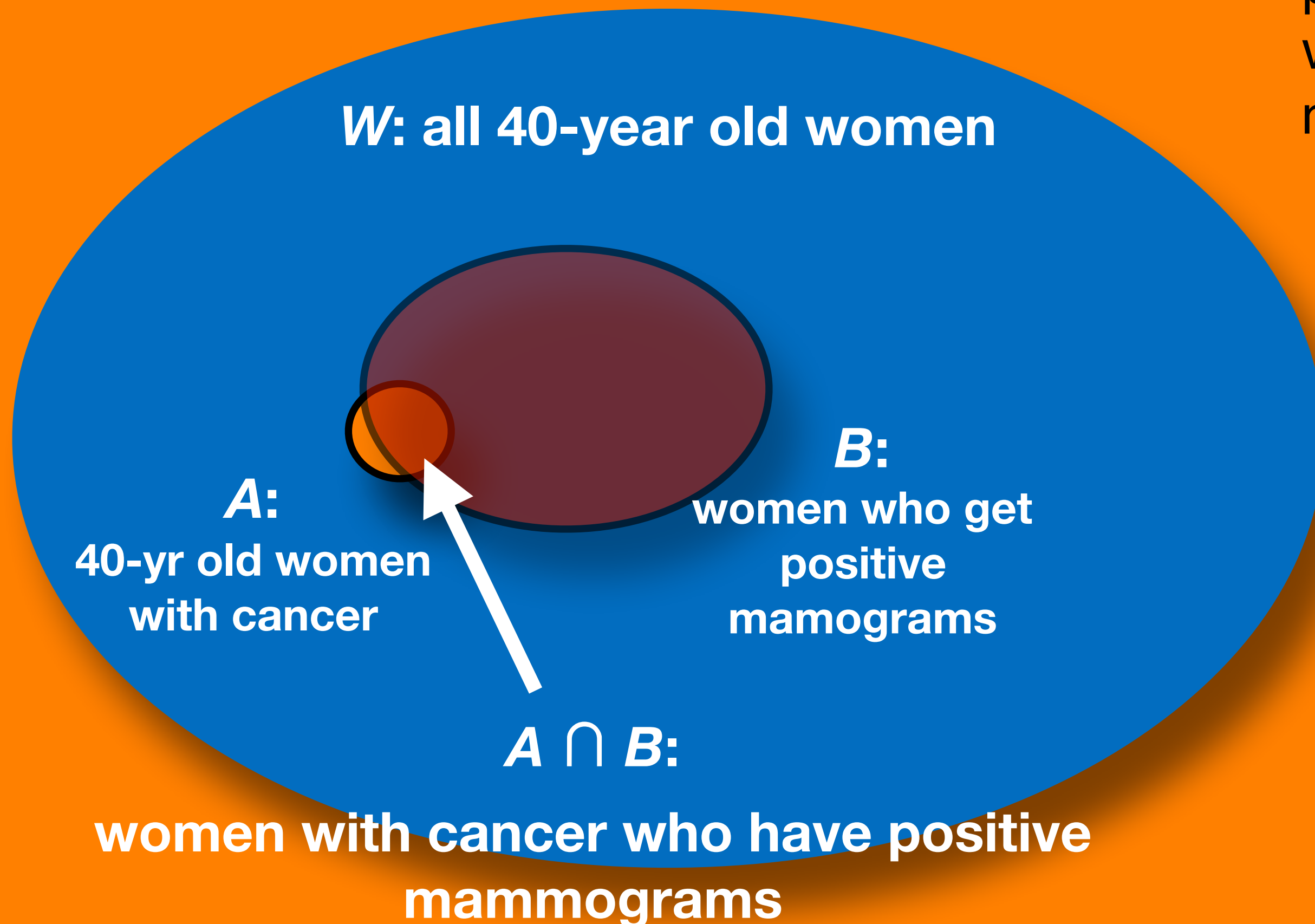
$$p(A \cap B) = \frac{|A \cap B|}{|W|}$$

given you've got cancer, what's the likelihood you'll also have a positive mammogram:

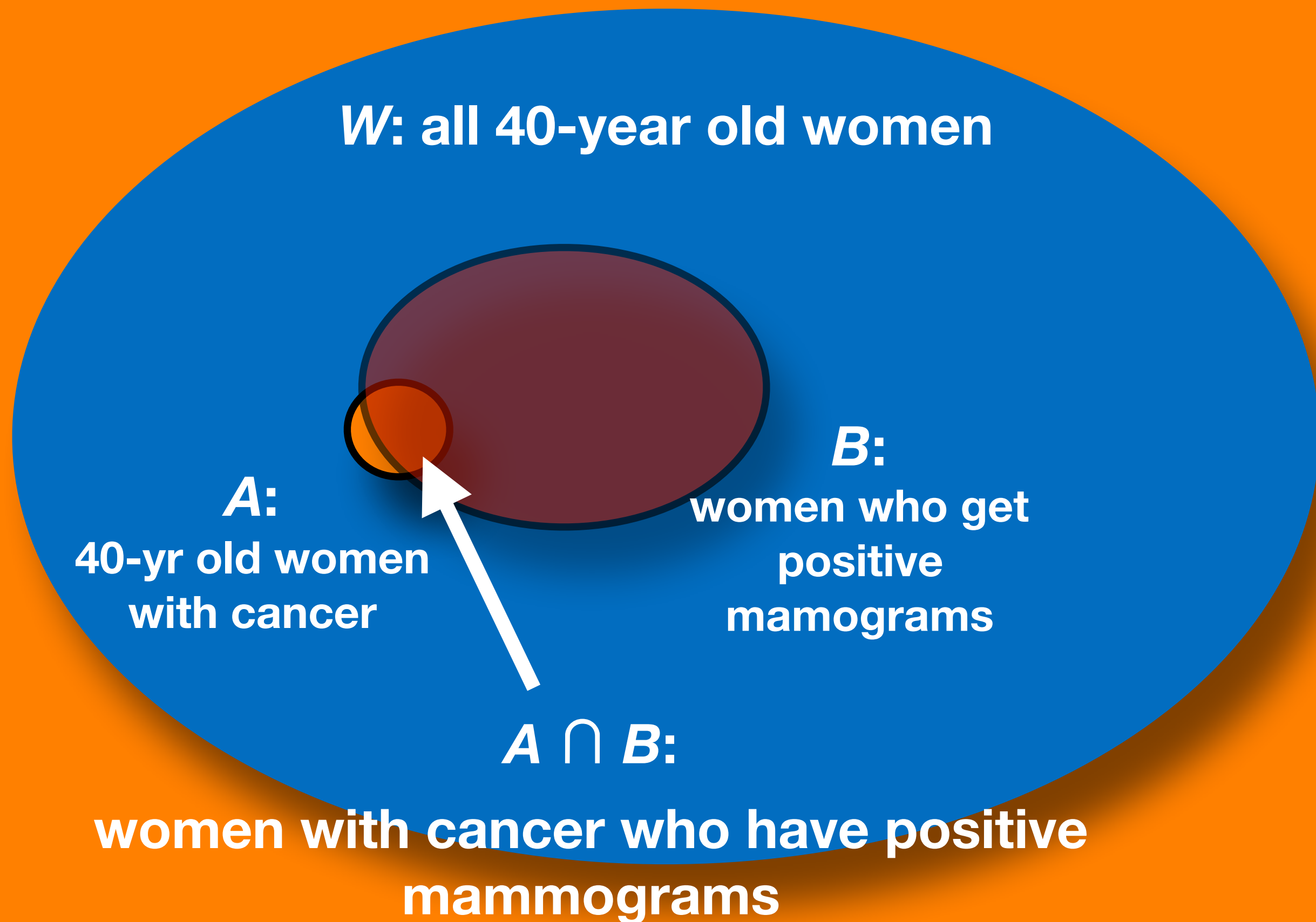
$$p(B|A) = \frac{|A \cap B|}{|A|} \times \frac{|W|}{|W|} = \frac{p(A \cap B)}{p(A)}$$

given you've had a positive mammogram, what's the probability you've got cancer:

$$p(A|B) = \frac{|A \cap B|}{|B|} \times \frac{|W|}{|W|} = \frac{p(A \cap B)}{p(B)}$$



- 1% of women at age forty have breast cancer.
- 80% of women with breast cancer have positive mammograms.
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- A woman in this age group had a positive mammogram in a routine screening. What is the probability that she has breast cancer?



$$p(B|A) = \frac{p(A \cap B)}{p(A)} \qquad p(A|B) = \frac{p(A \cap B)}{p(B)}$$

$$p(A|B) = \frac{p(B|A)p(A)}{P(B)}$$

$$p(B|A) = 0.8$$

$$p(A) = 0.01$$

$$p(B) = 0.8p(A) + 0.096(1 - p(A)) \approx 0.1$$

$$\Rightarrow p(A|B) \approx 0.078$$

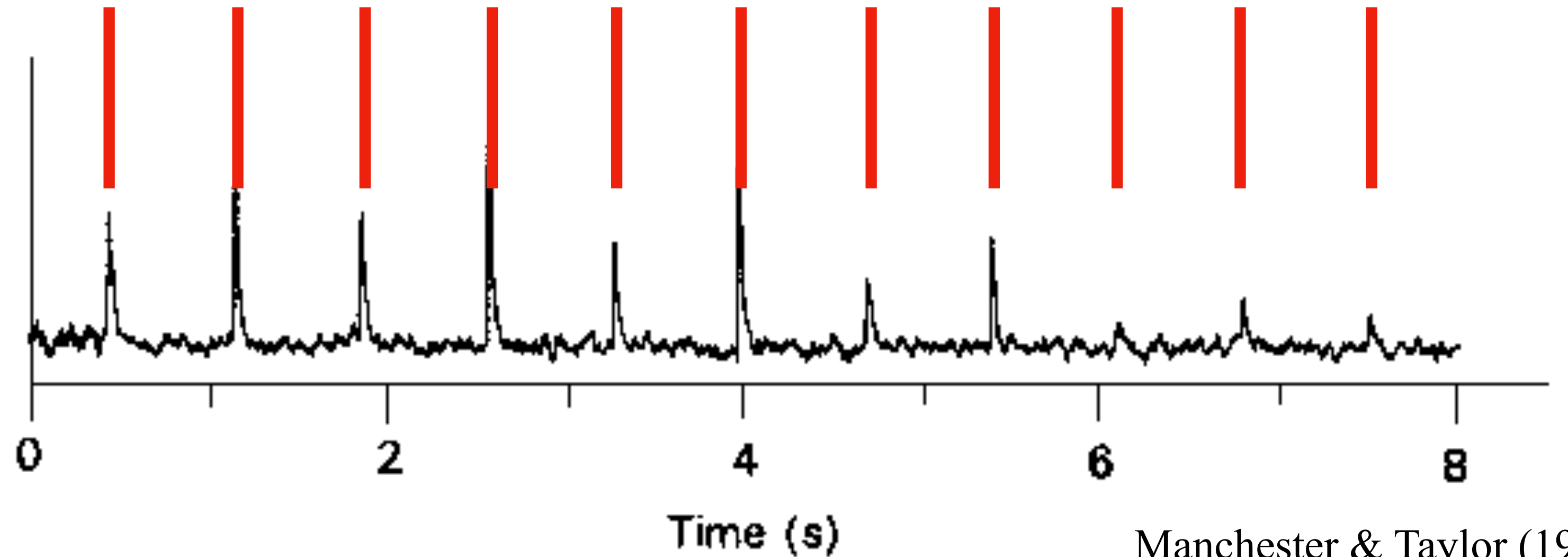
If you're a 40-year old woman, and have a positive mammogram, there's only a ~7.8% chance you have cancer!

$$p(A|B) = \frac{p(B|A)p(A)}{P(B)}$$

What about pulsar astronomy?

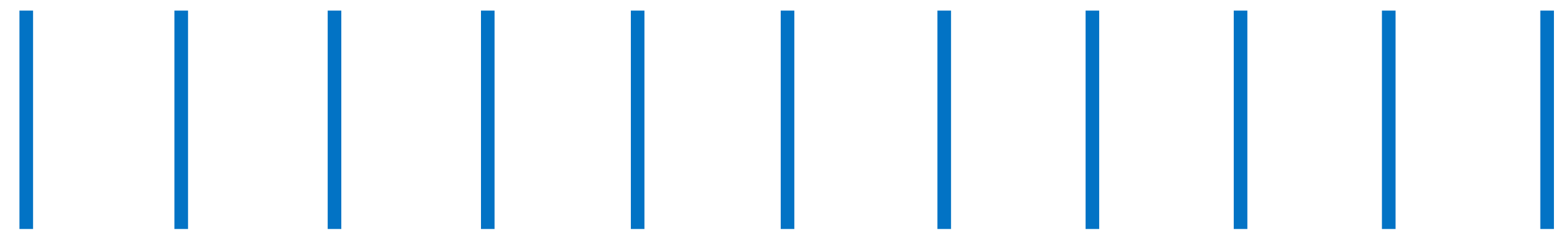
data: d

(pulse times of arrival)



Model: M

with parameters: θ



e.g., pulsar spin frequency, f

We want to calculate the probability of certain model parameters given the data: $p(\vec{\theta} | d)$

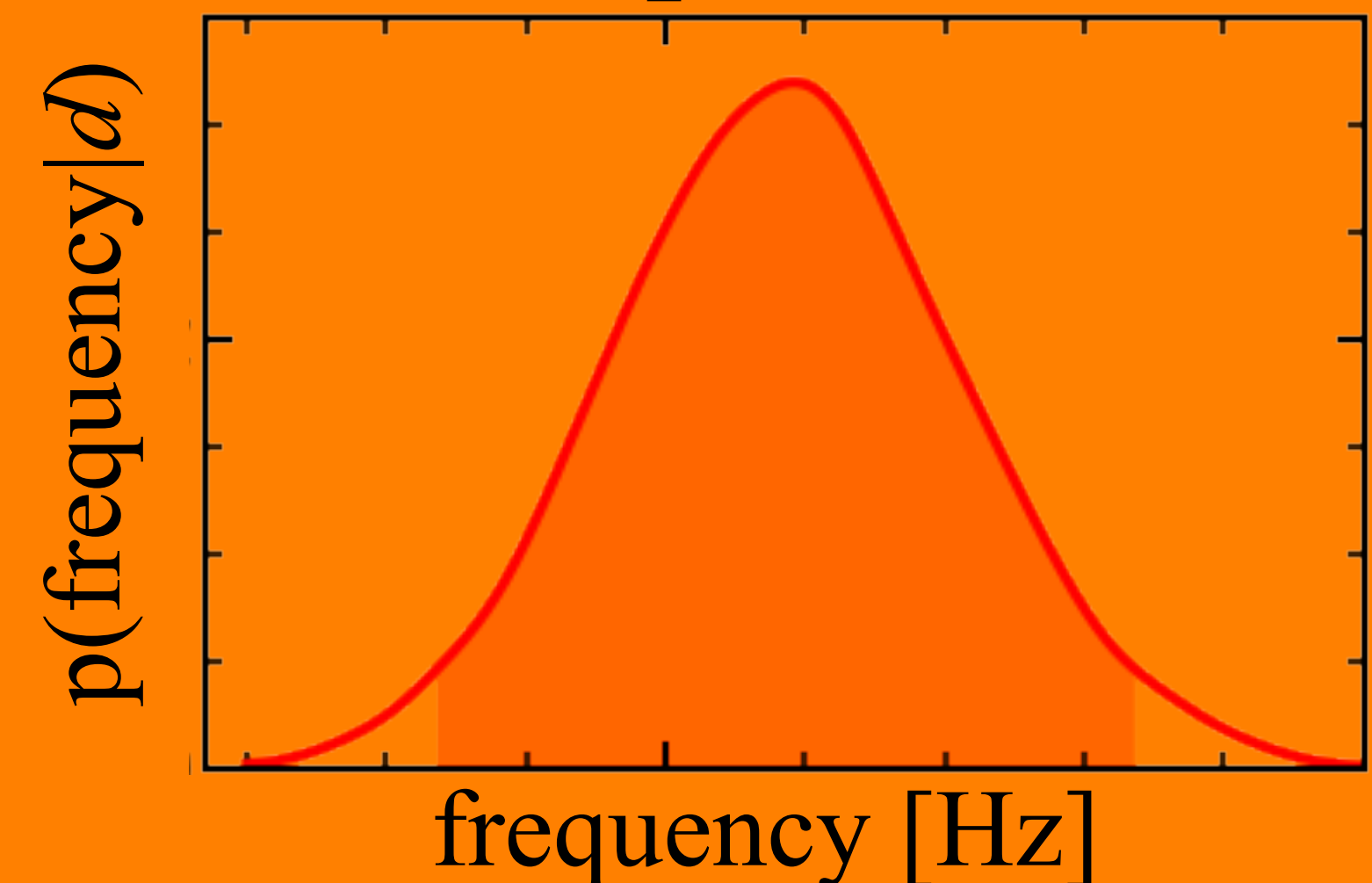
Likelihood: probability of the data, given the model.

Prior: our prior knowledge of the parameters of the model

$$p(\theta|d) = \frac{L(d|\theta)\pi(\theta)}{Z(d)}$$

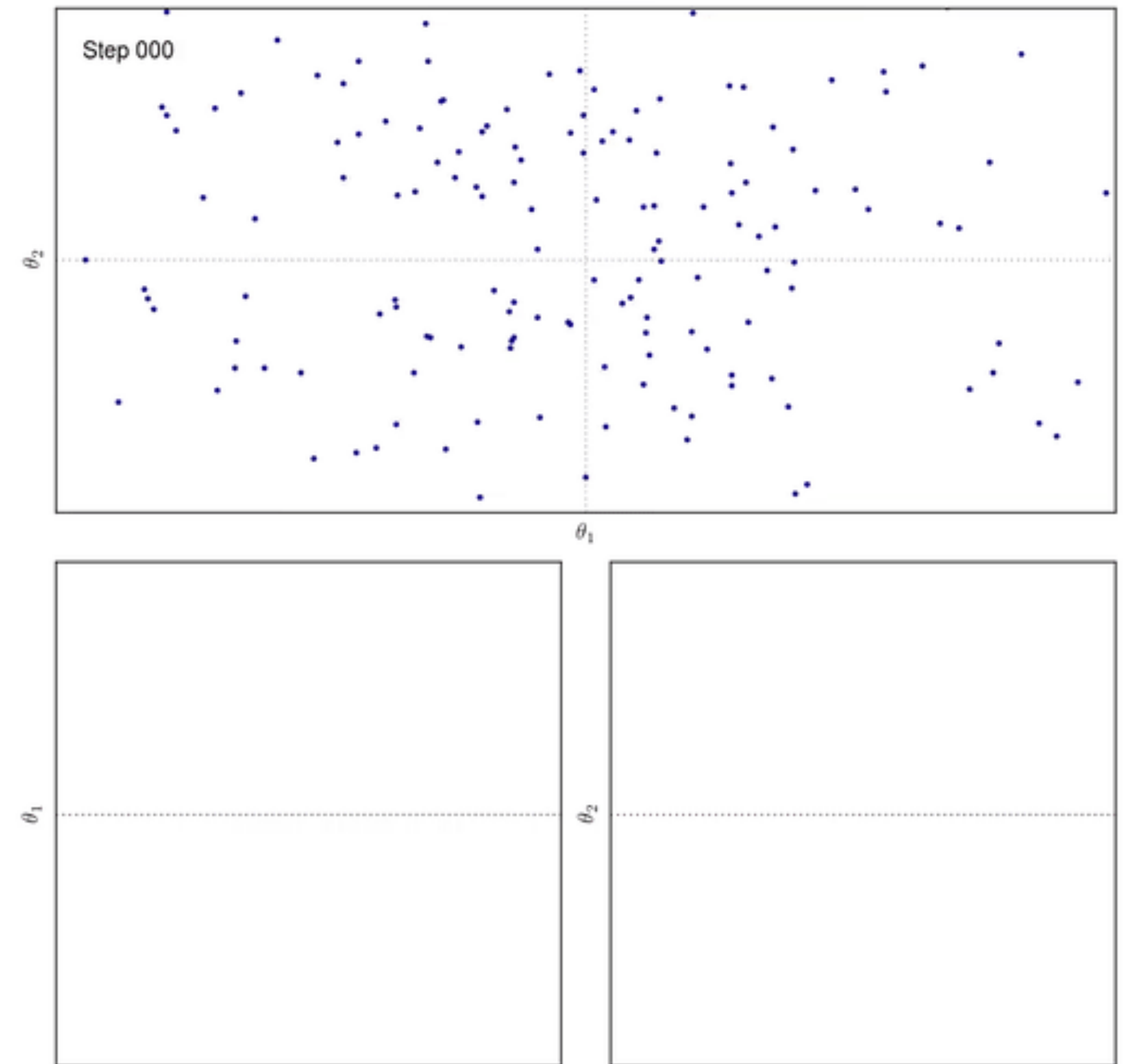
The game: Evaluate this equation for all values of the model parameters

Evidence: (only depends on the data, for parameter estimation is just a normalisation constant; let's get back to this)

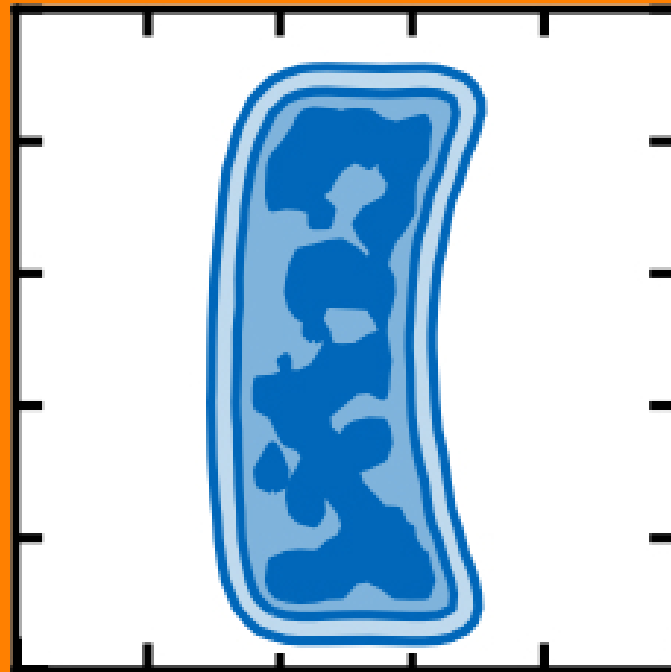


in reality, for large dimensional spaces, direct integration is computationally infeasible

- Markov chain Monte Carlo
- Nested sampling
- etc



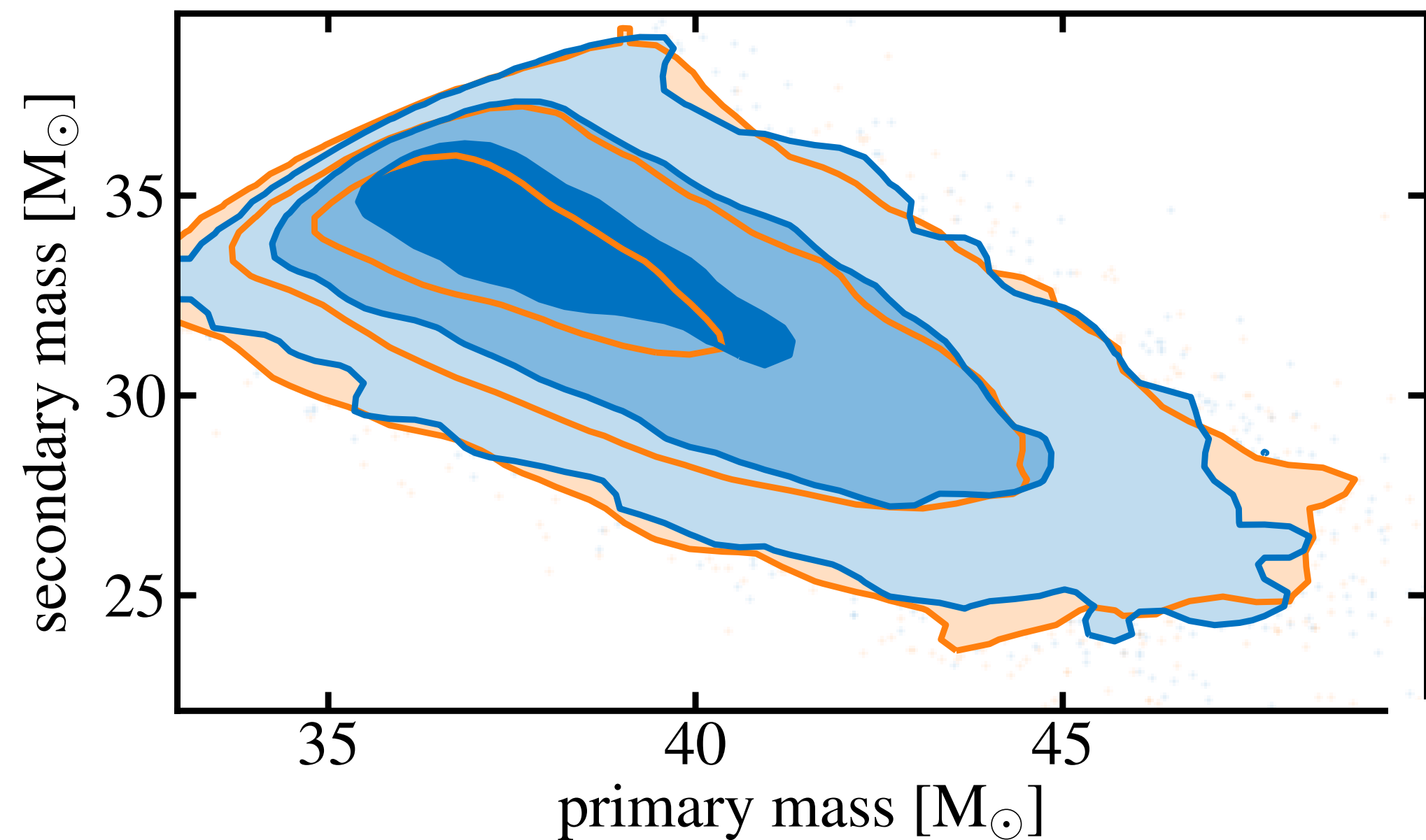
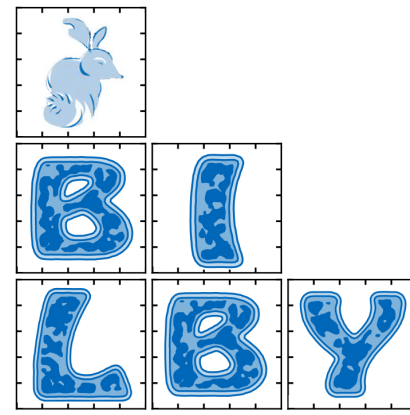
The user-friendly Bayesian inference library



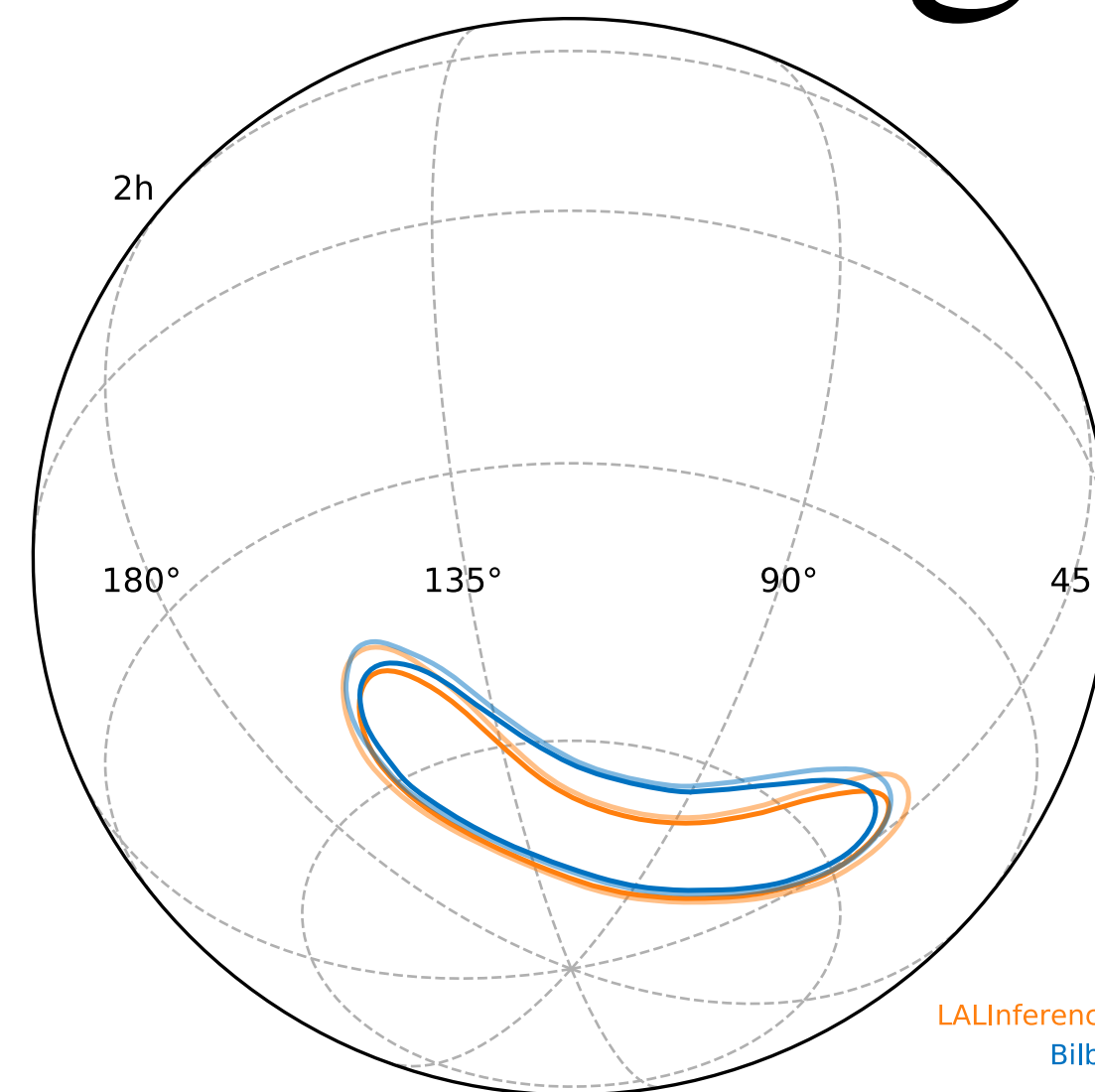
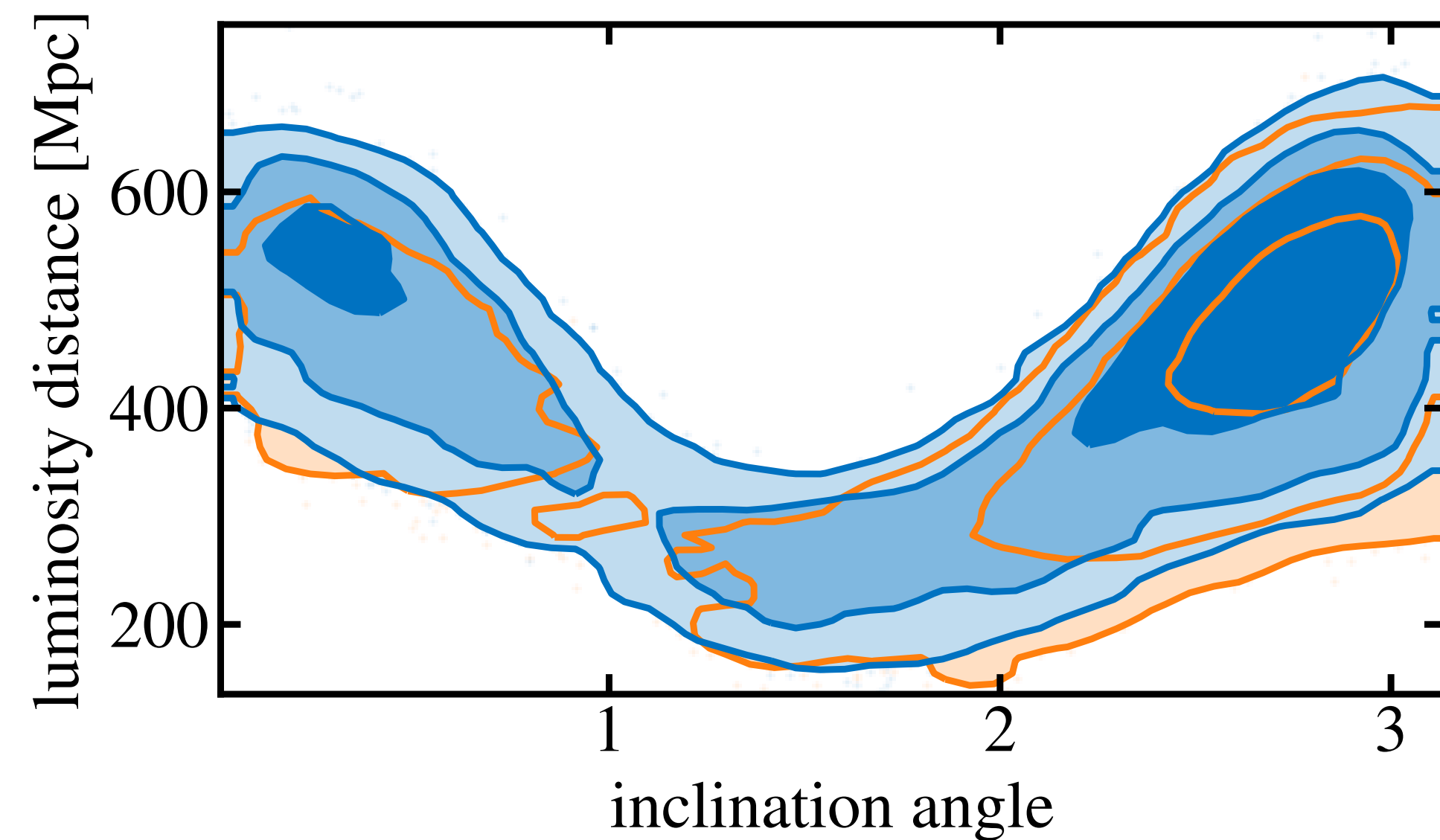
A versatile parameter-estimation code being adopted for
production science in next LIGO observing run

git.ligo.org/lscsoft/bilby/

Ashton, Hübner, PL, Talbot + (2019)

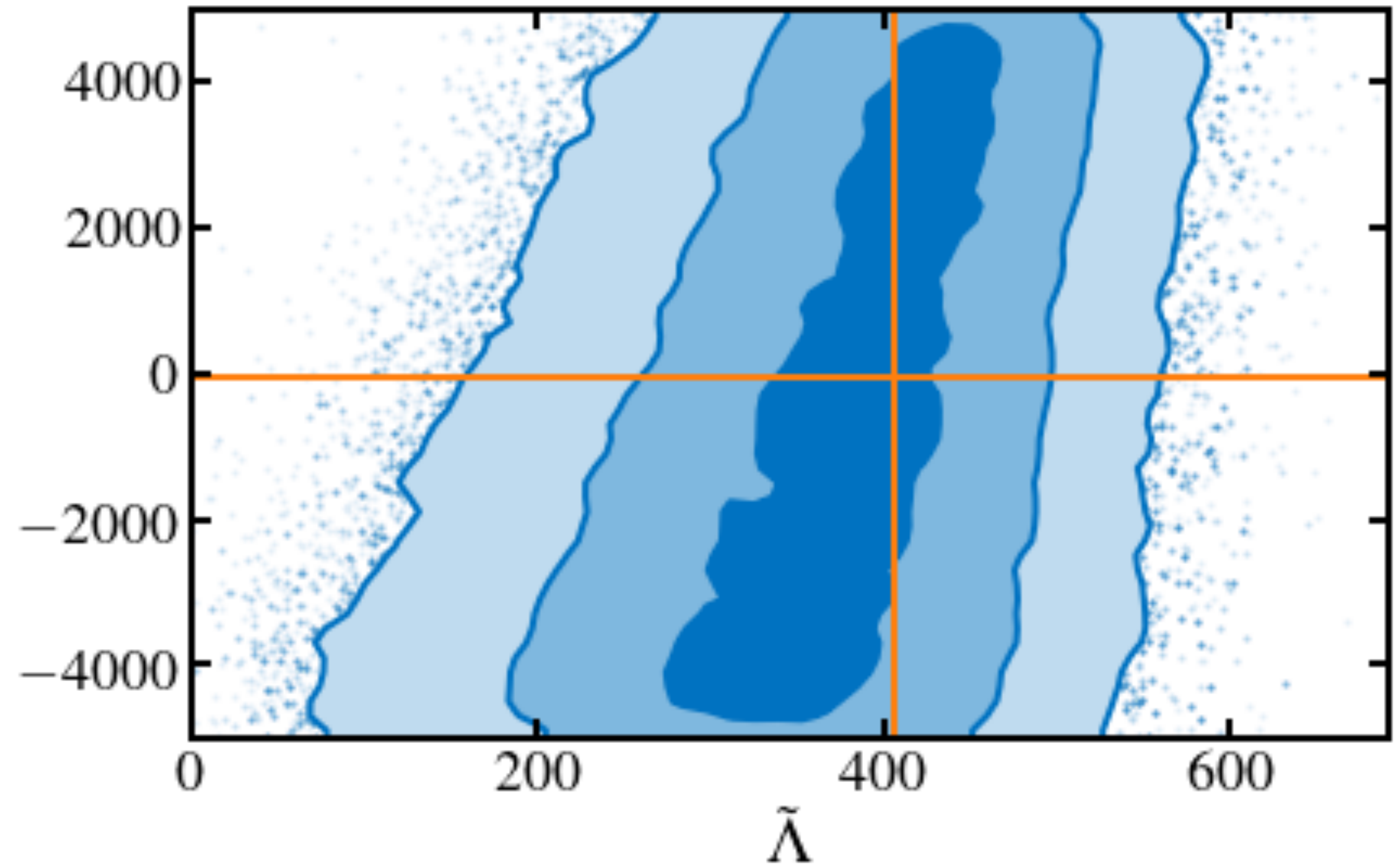
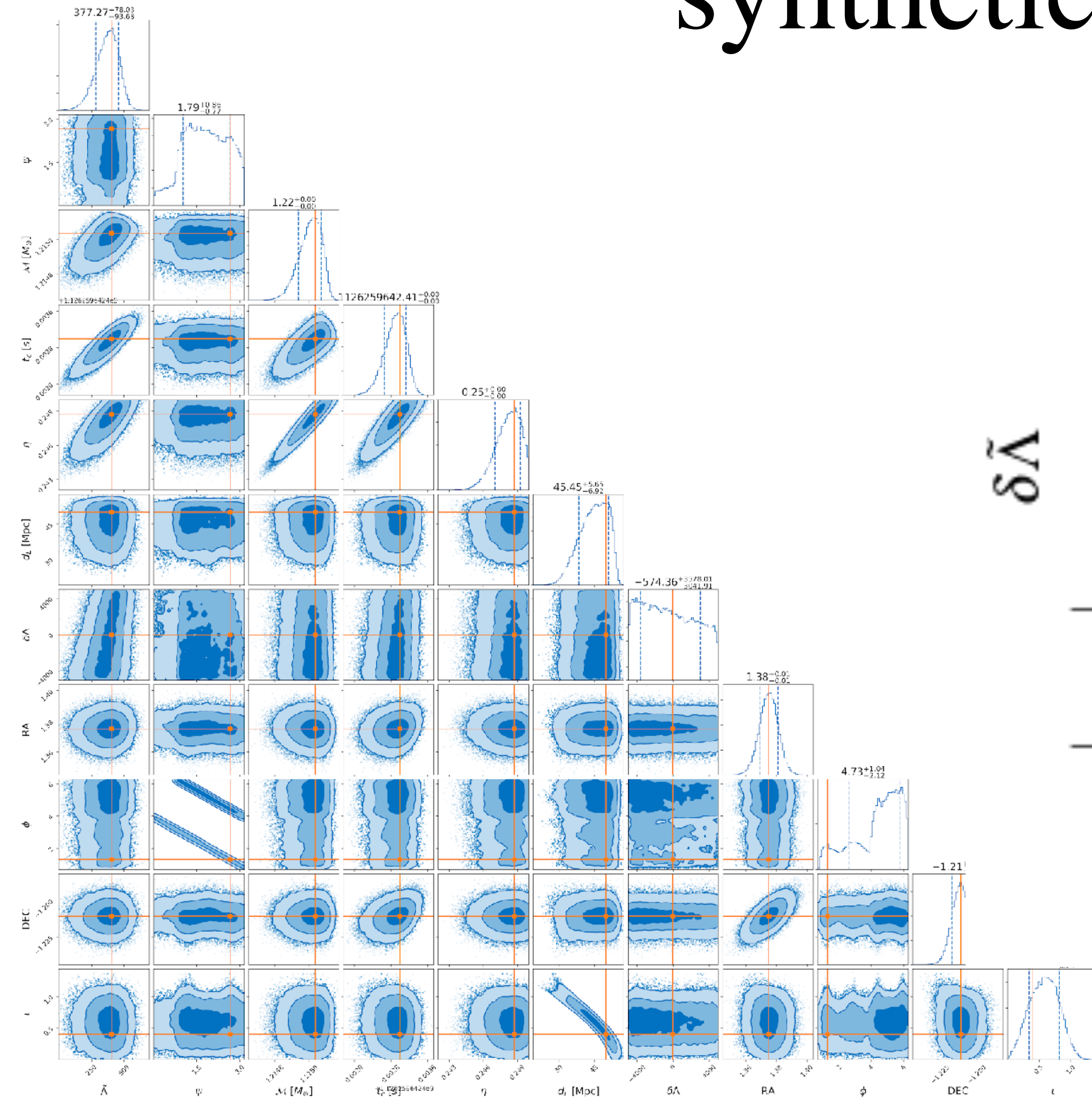


ground-based gravitational-wave experiments



Ashton, Hübner, PL, Talbot + (2019)

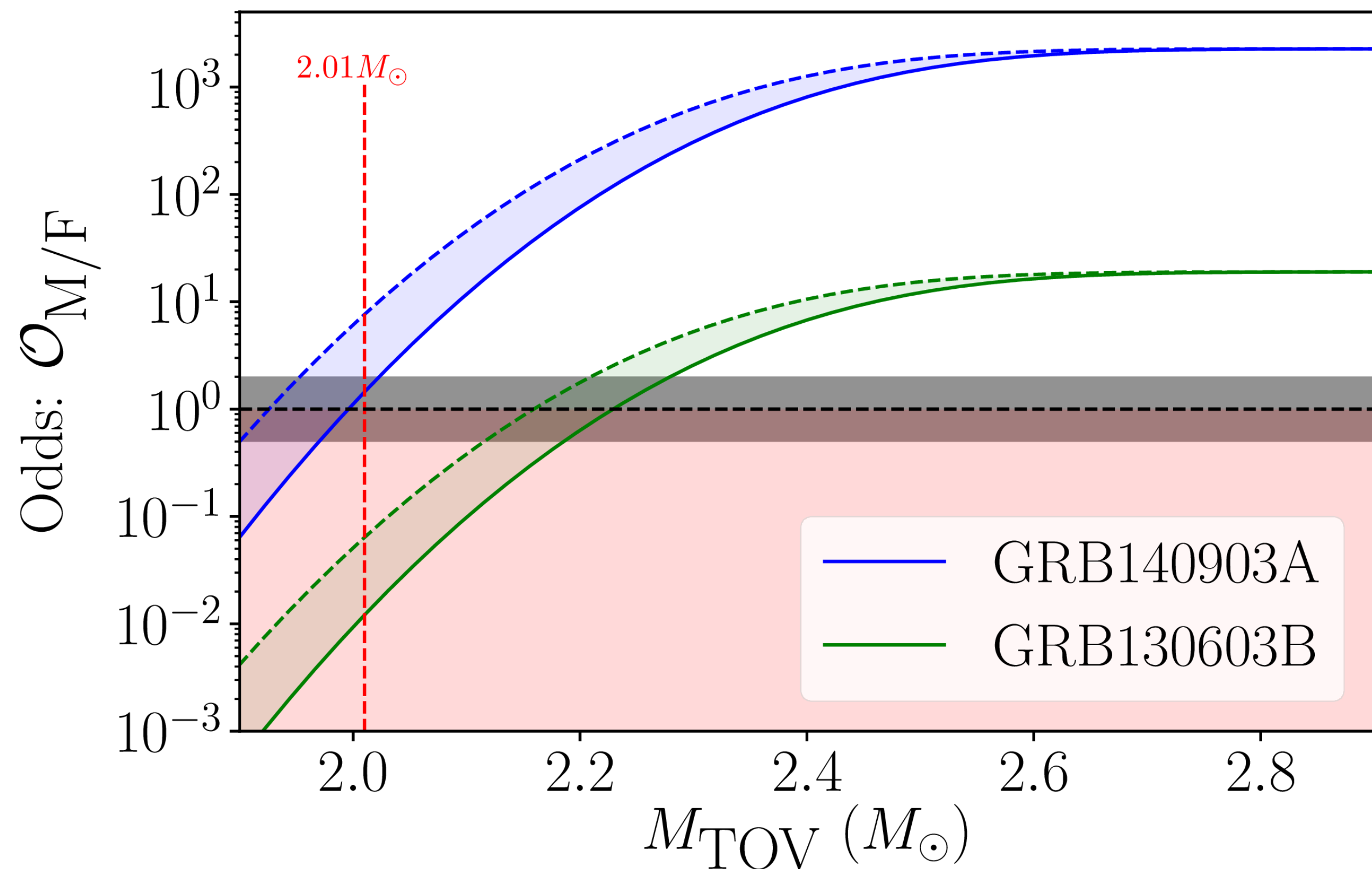
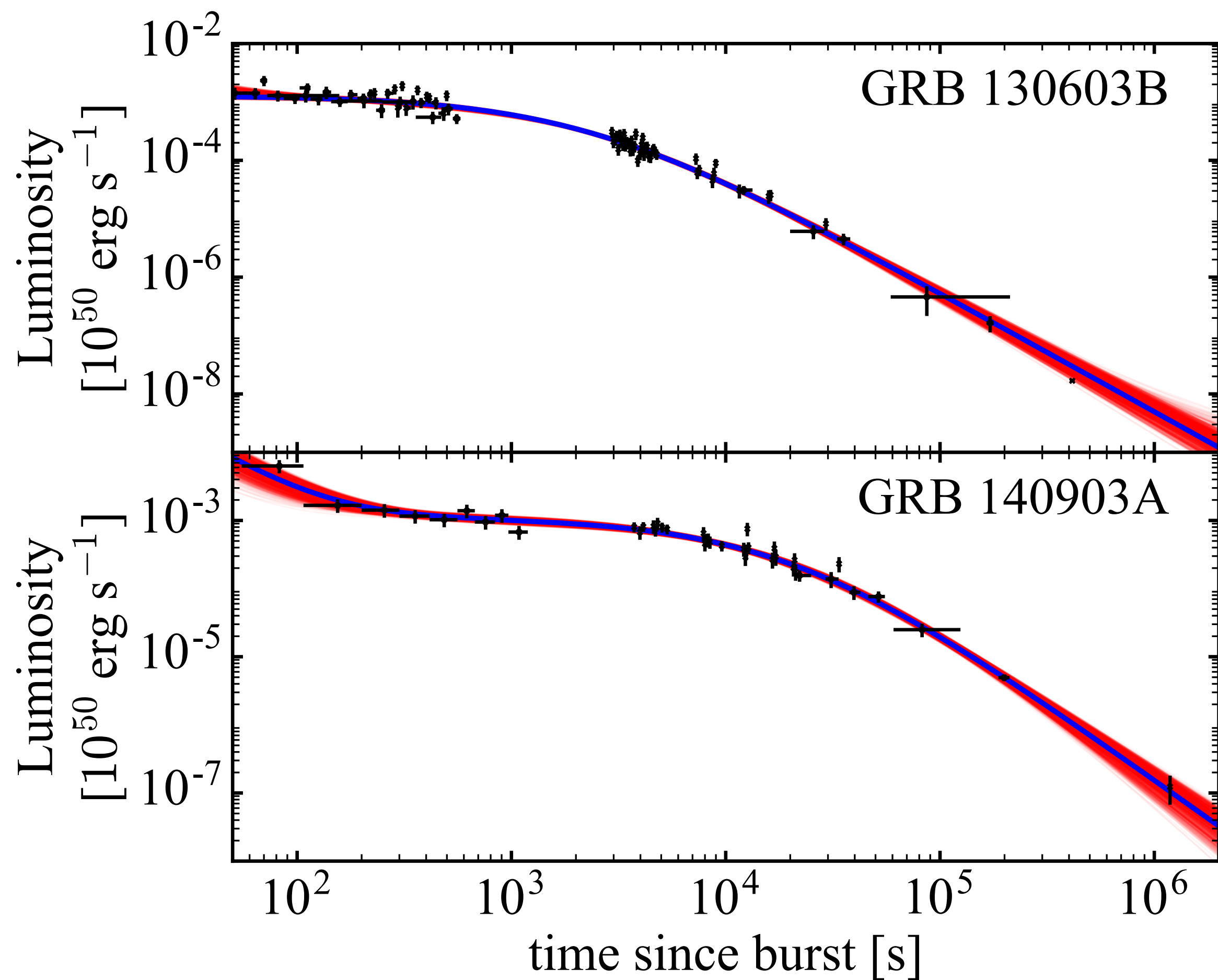
synthetic neutron stars



Ashton, Hübner, PL, Talbot + (2019)

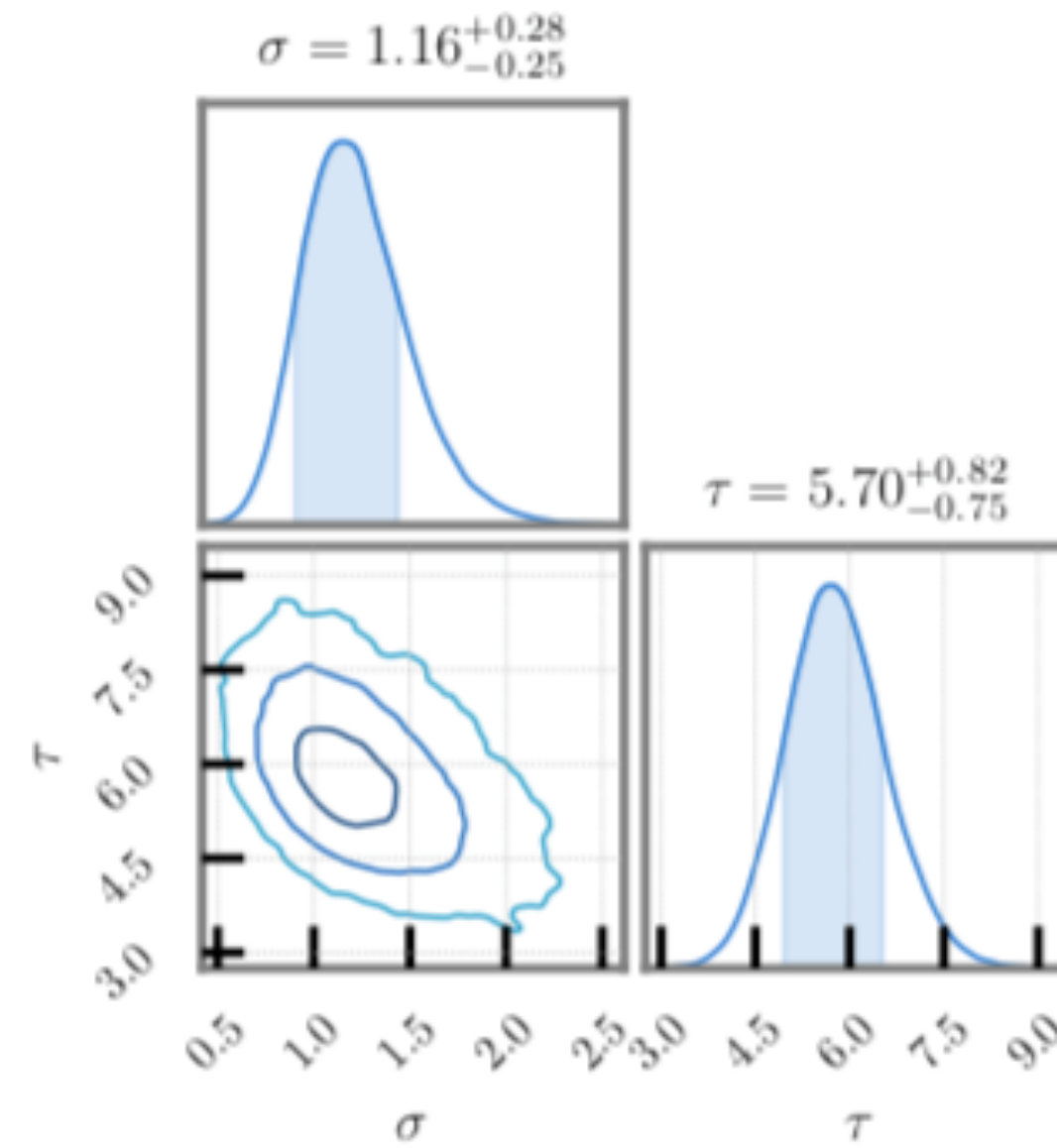
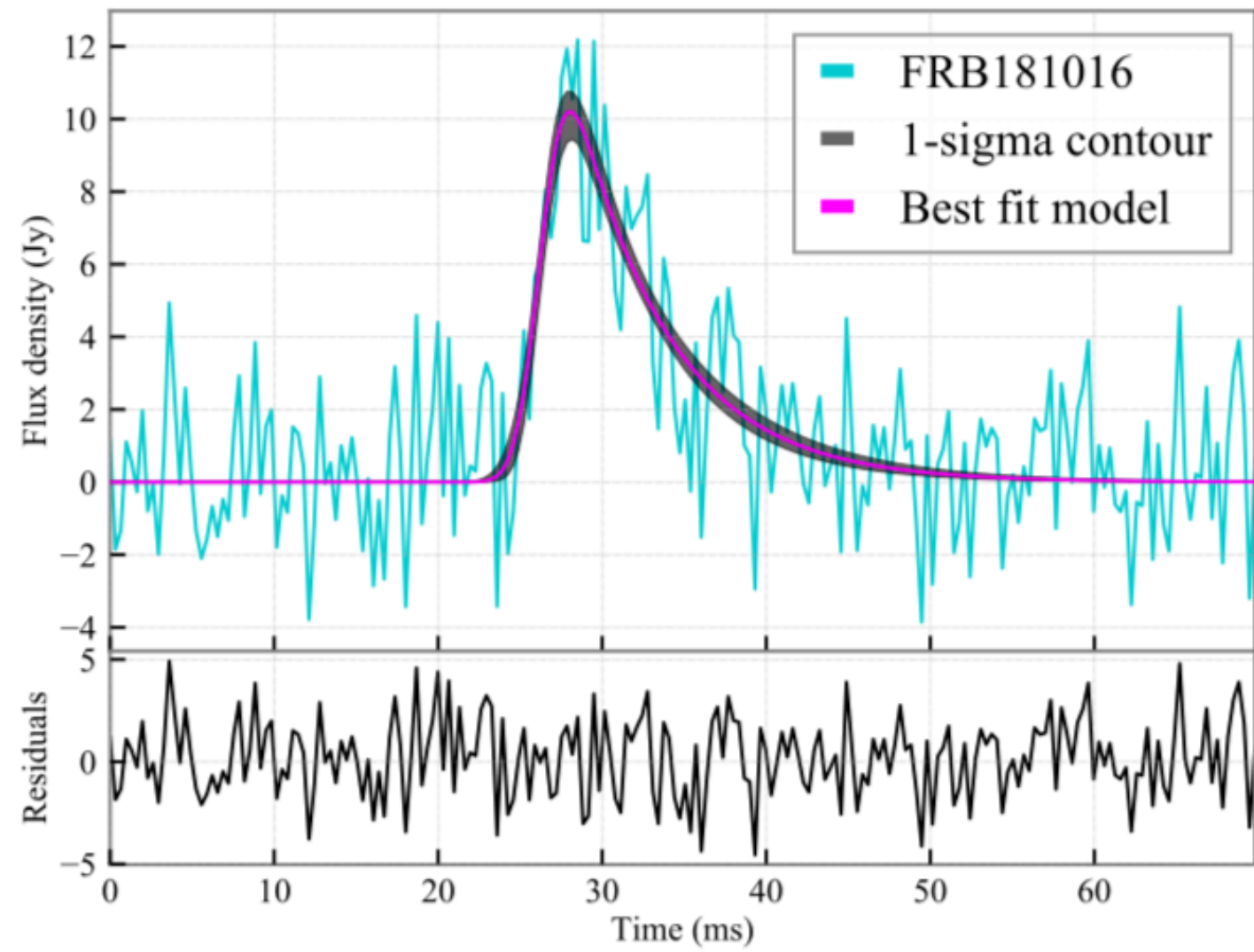
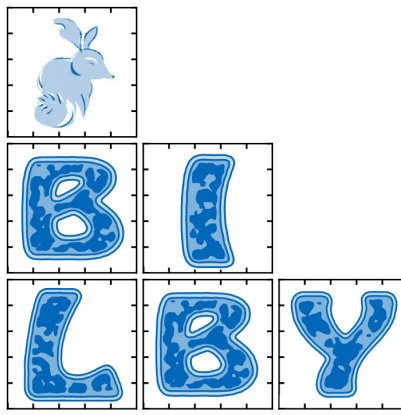
x-ray light curves of gamma-ray bursts

do millisecond magnetars exist?

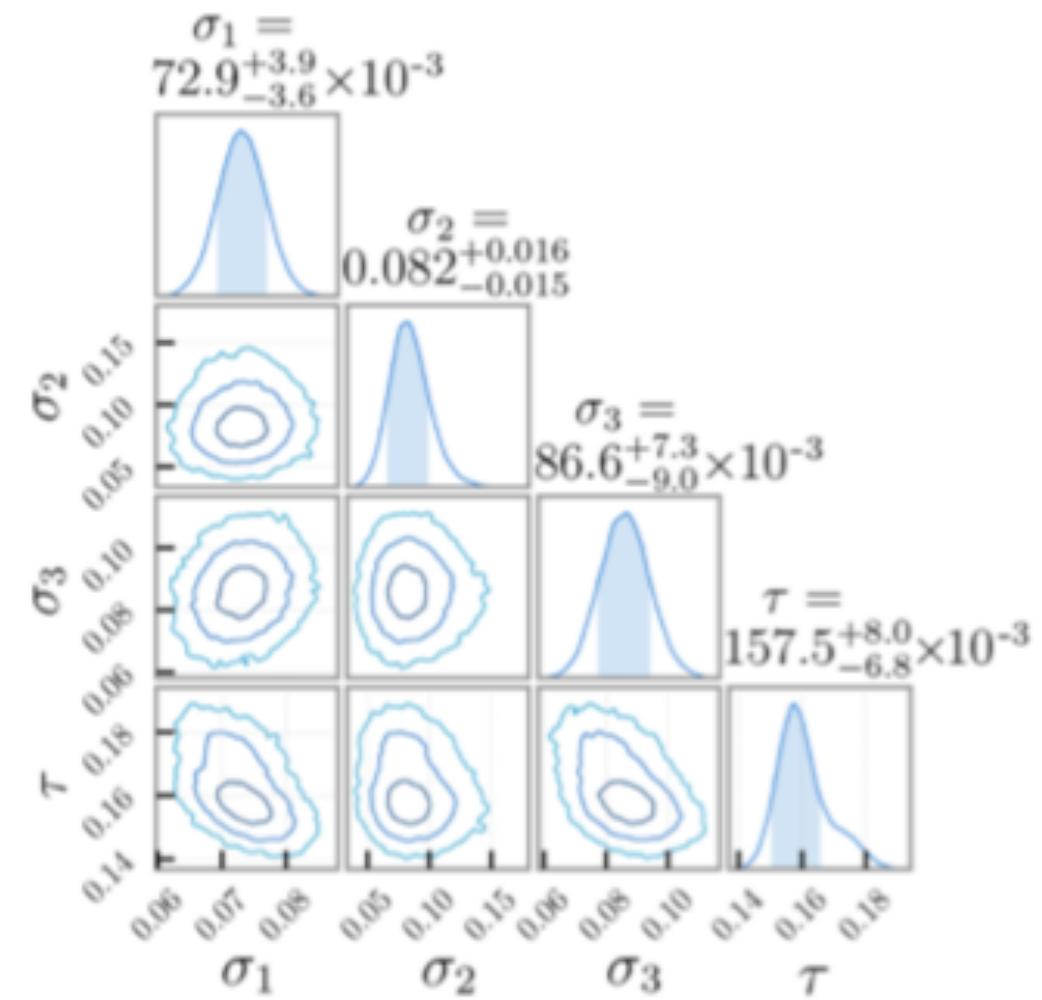
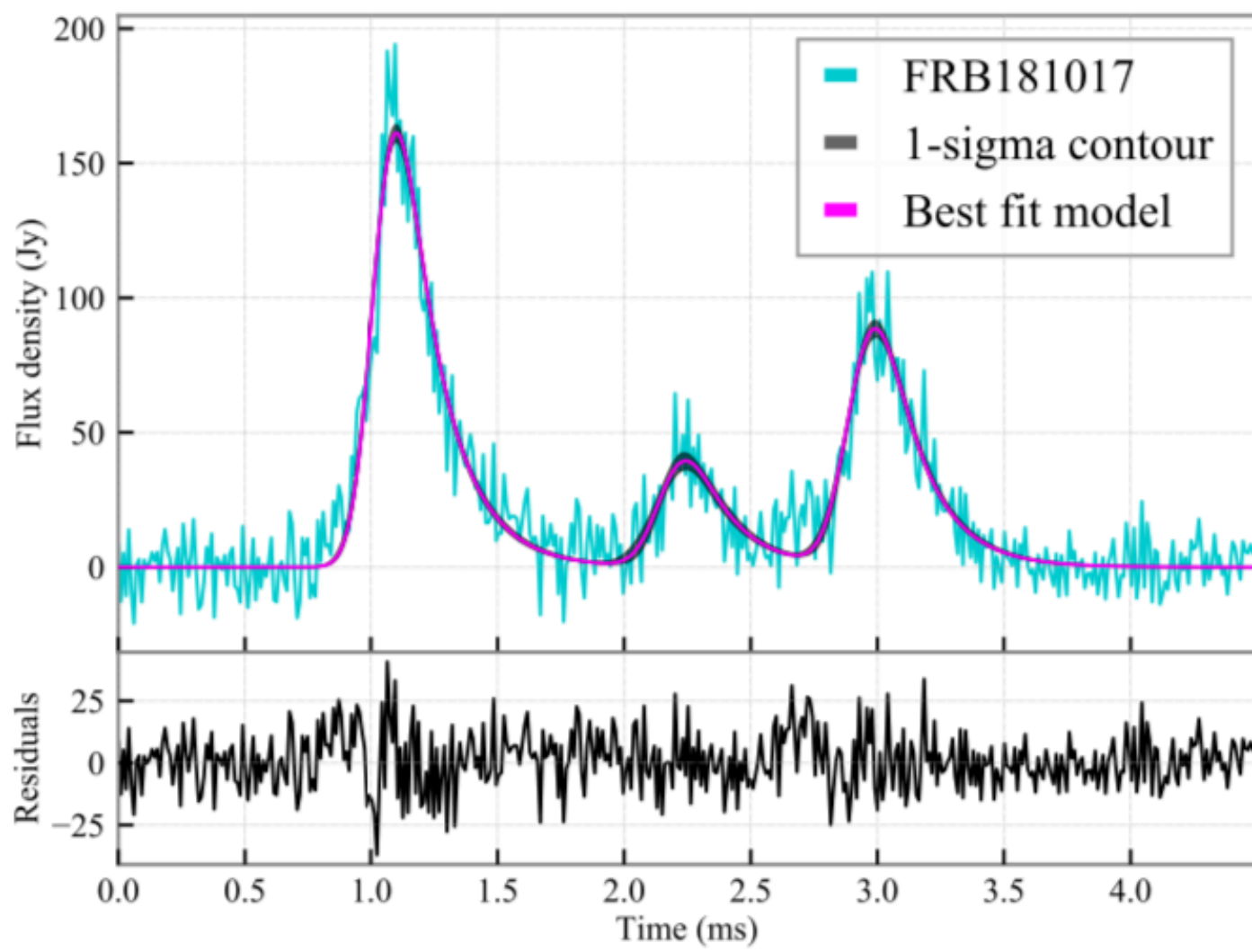


PL, Leris, Rowlinson & Glampedakis (2017)

Sarin, PL, Ashton (2019)

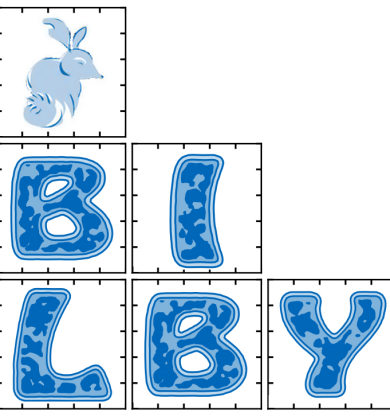


Fast Radio Bursts

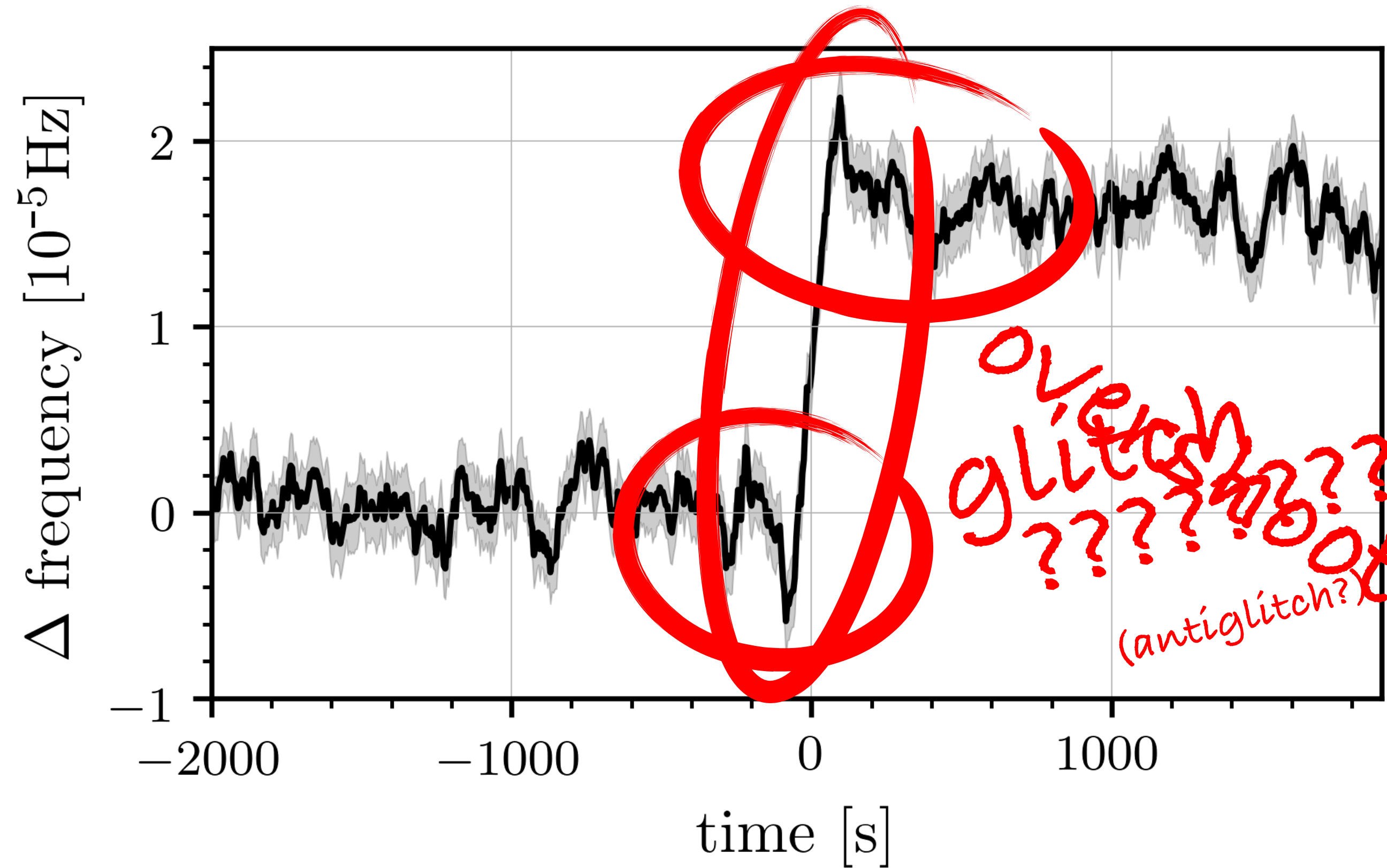


Farah et al. 2019

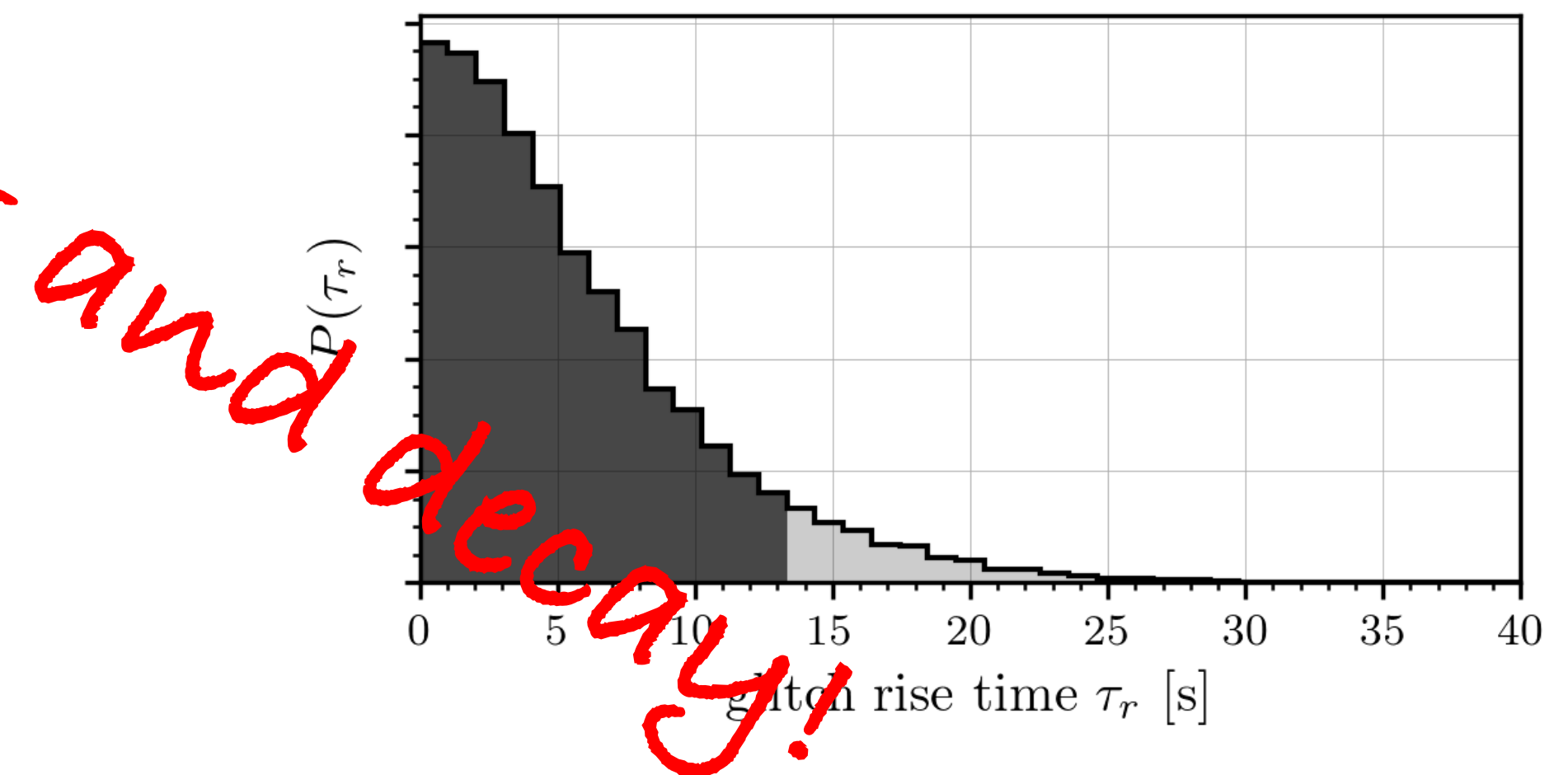
finally back to pulsars!



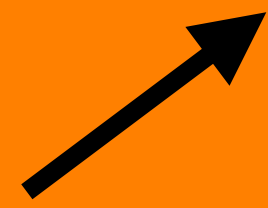
The 2016-vela glitch



- 200-s sliding windows
- each window, Bayesian inference using a constant-frequency model
- fit timing model with basic glitch, derive posterior for the rise time



$$p(\theta|d) = \frac{L(d|\theta)\pi(\theta)}{Z(d)}$$



Evidence: used for doing model selection!

$$\int d\theta p(\theta|d) = 1 \quad \longrightarrow \quad Z(d) = \int d\theta L(d|\theta)\pi(\theta)$$

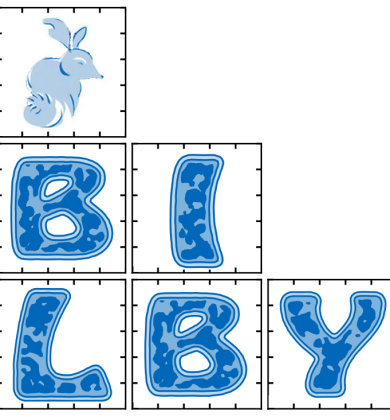
Given two models: $M_1(\theta_1)$ and $M_2(\theta_2)$, the Bayes factor is defined as the ratio of the two evidences:

$$BF = \frac{Z_1}{Z_2} = \frac{\int d\theta_1 L(d|\theta_1)\pi(\theta_1)}{\int d\theta_2 L(d|\theta_2)\pi(\theta_2)}$$

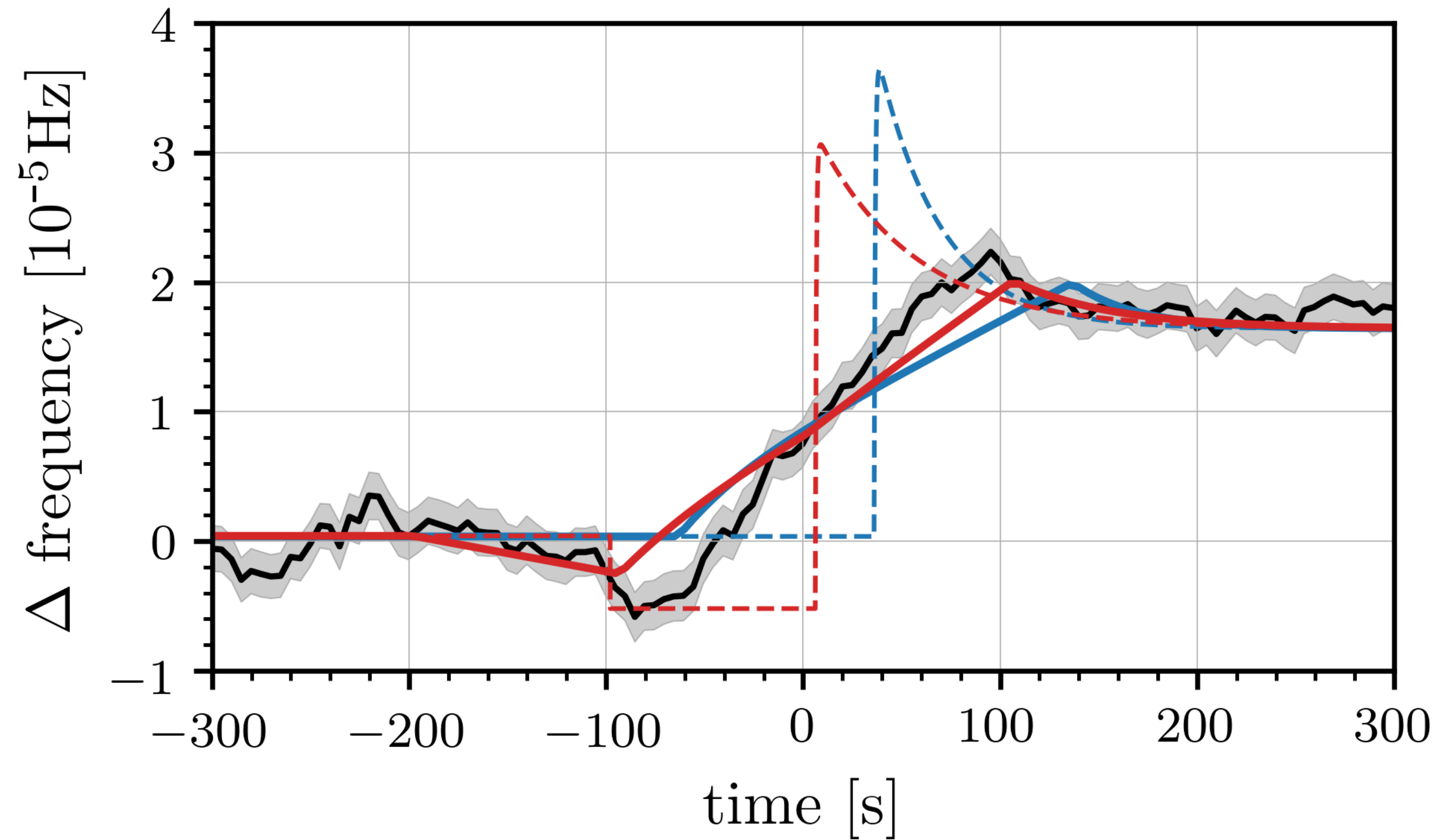
The Bayes factor (almost*) gives the odds between any two models. e.g.,
model 1 is preferred over model 2 with XX Bayes factor...

* technically, it's the "Odds", which multiplies the Bayes factor by the prior odds. But for simple situations, the prior odds equal one!

Bayesian model selection



The 2016-vela glitch



- does the overshoot-decay exist?

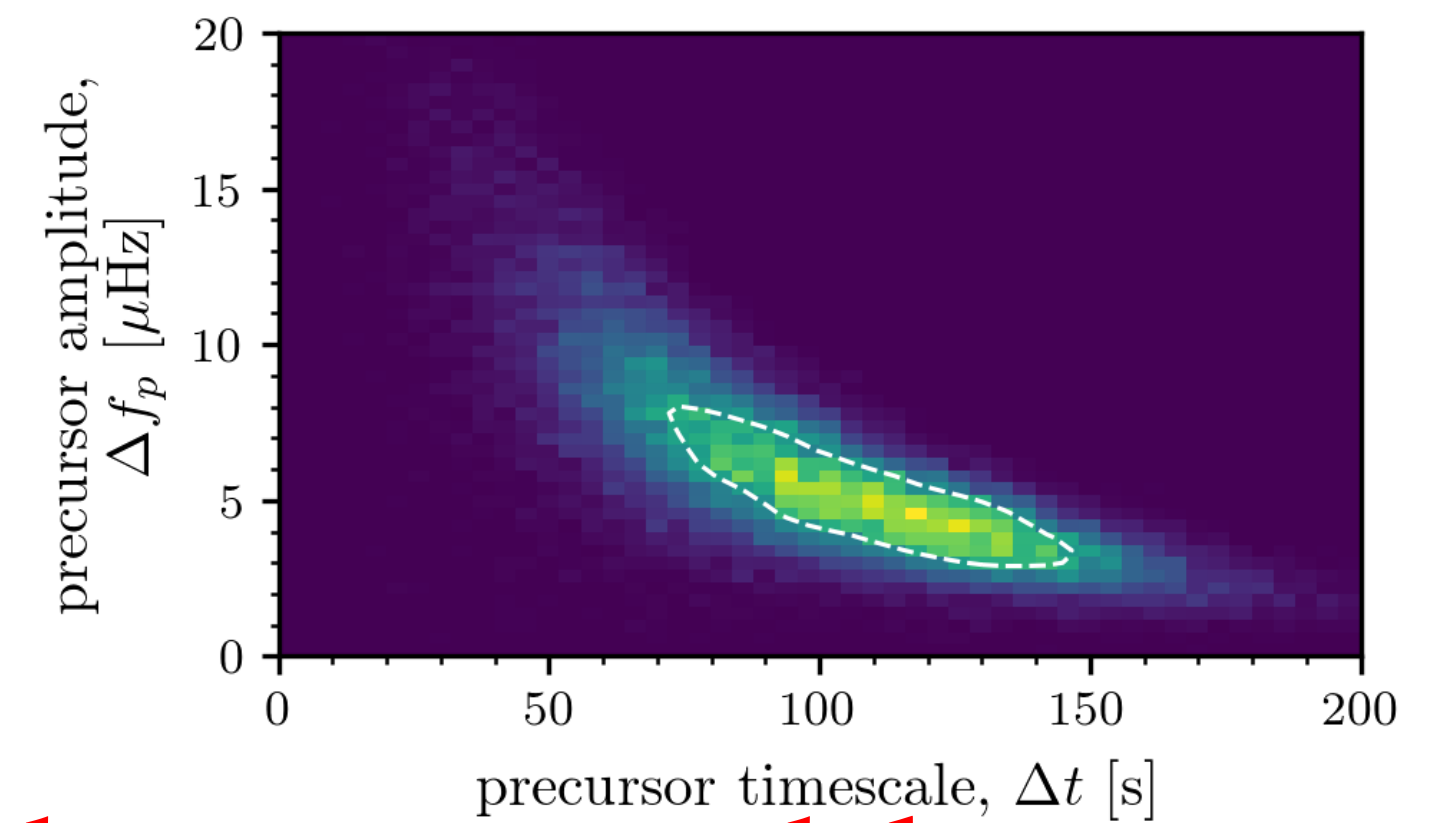
yes! BF ~ 125

(over simple step glitch)

- does the slow-down occur?

probably... BF ~ 3

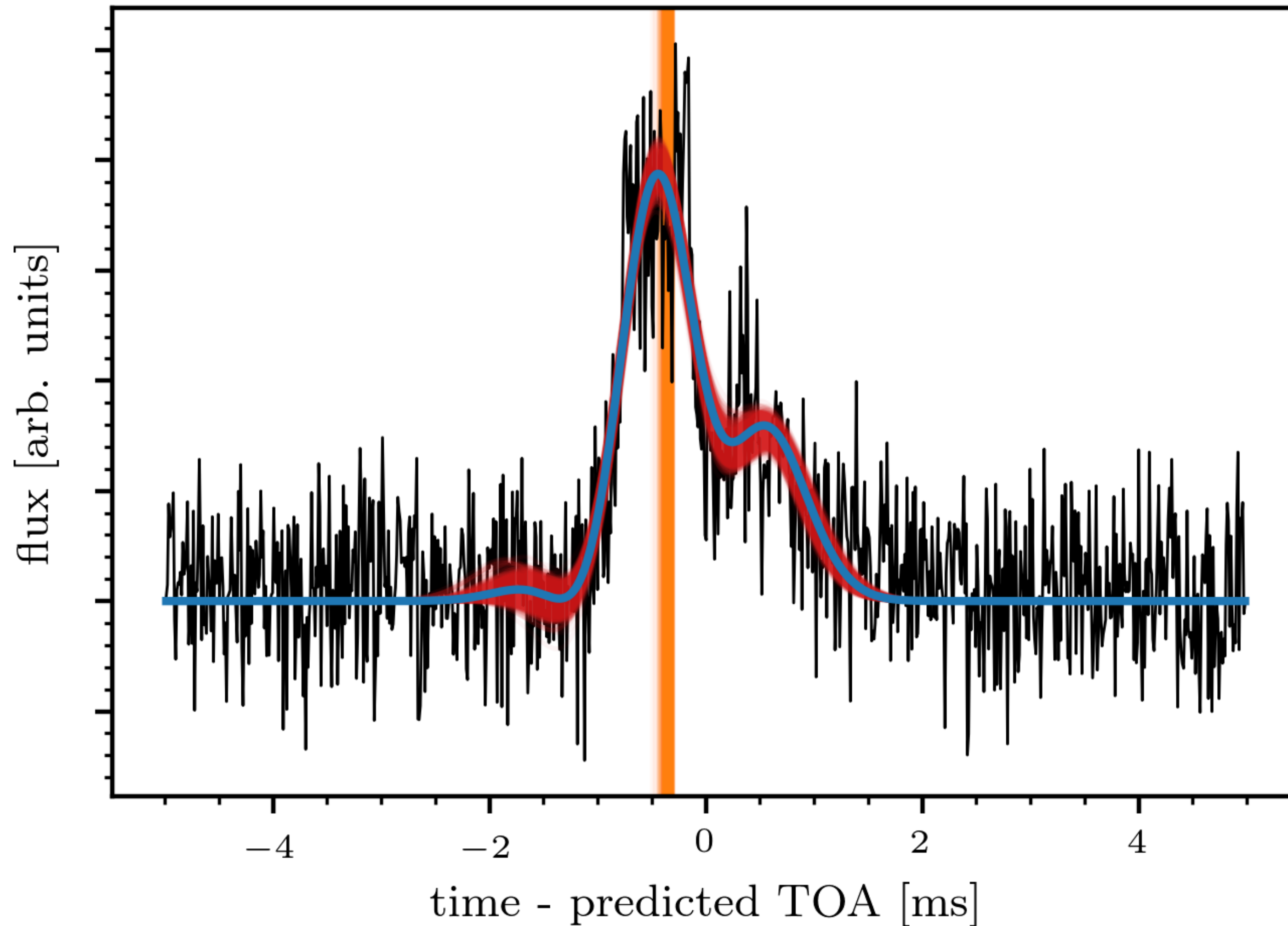
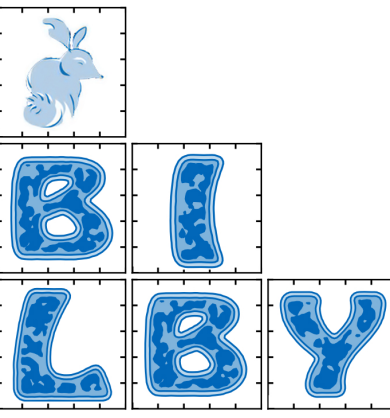
(over overshoot-decay-only model)



word of caution: you're only as good as your models...

Ashton, PL, Graber, Palfreyman (submitted)

pulse-profile modelling



Greg Ashton has developed this;
speaking about it tomorrow!

- Pulsar timing uses pulse time-of-arrivals
- this is not the right way to do things
- Reverend Thomas Bayes would:
 - do full parameter estimation for each pulse
 - use a hierarchical model for the spin evolution, etc., of the pulsar
- hierarchical model includes things like
 - gravitational-wave background
 - pulsar red- and white-noise parameters
 - solar-system parameters
 - etc

plot courtesy of Greg Ashton, via pyglit/Bilby

**“Bayesian inference is the
future of gravitational-wave
astronomy”**

Matilda B. Bilby*
*still not a real Bilby

