

Pulsars @ OzGrav

Highlights, challenges and opportunities
Ryan Shannon & Hannah Middleton, co-chairs





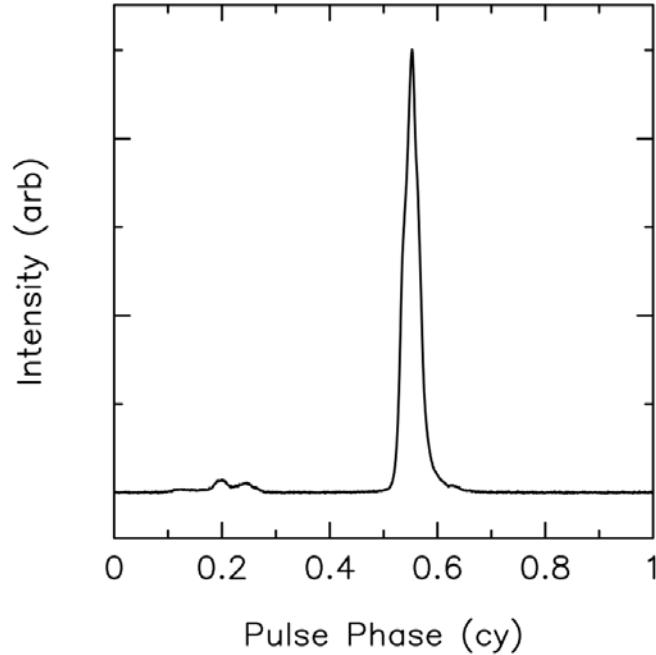
ARC Centre of Excellence for Gravitational Wave Discovery

#pulsar-inference (ozgrav slack)
#ppta (ozgrav slack)
meertime.slack.com
PPTA telecon
Australasian monthly pulsar telecon
- Clash with Melbourne astronomy seminar
IPTA-Gravitational Wave Analysis (East) telecon

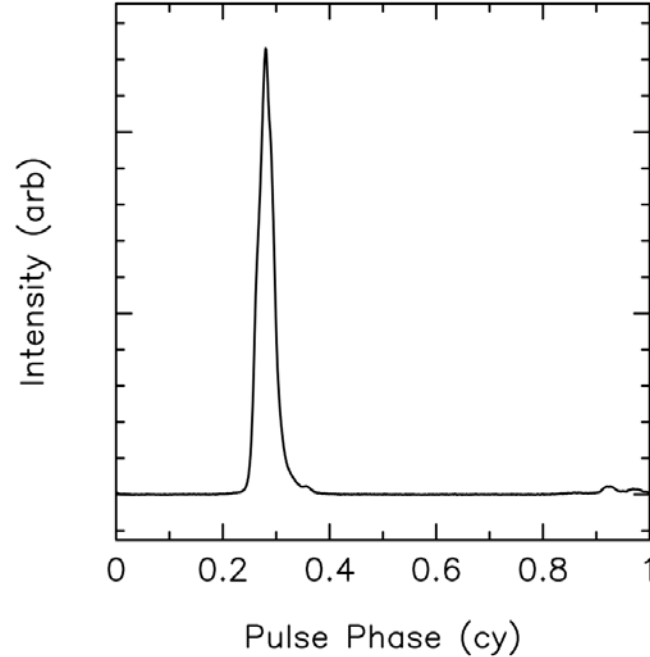


MeerKAT: A powerful pulsar telescope

5 days of PKS

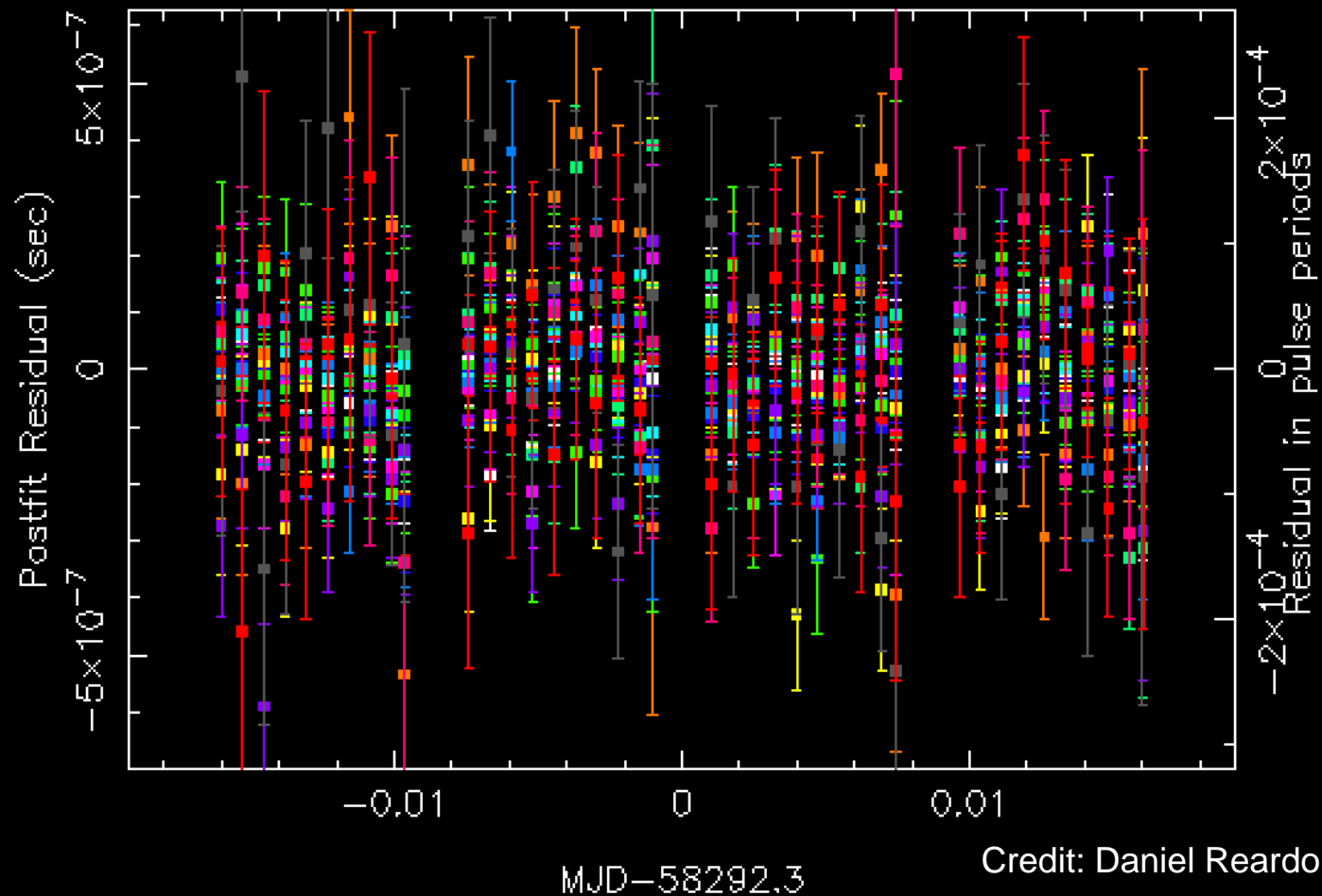


30 minutes of MK

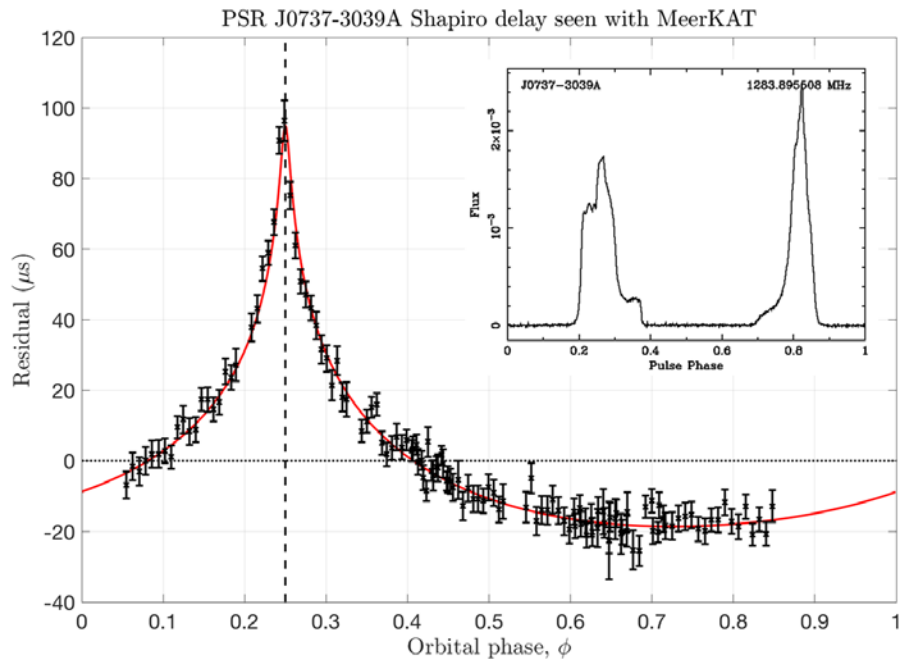


w tim Restart J2241-5236 (Wrms = 0.108 μ s) post-fit

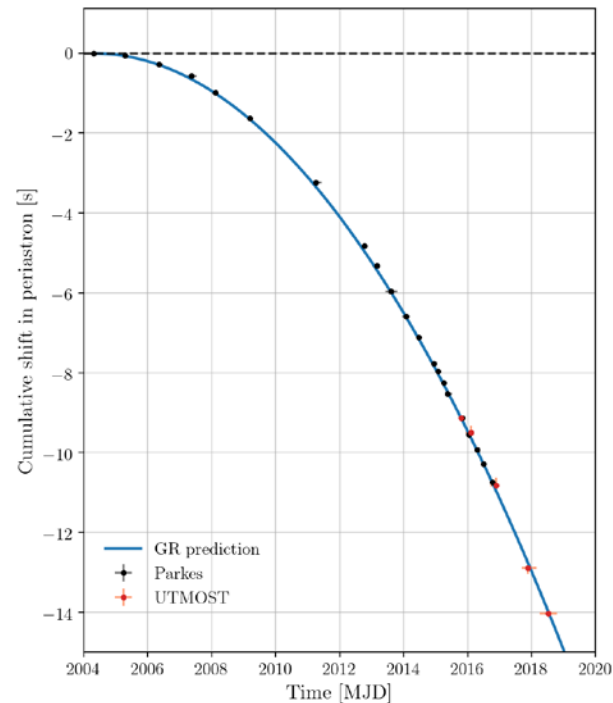
PSR J2241-5236
108 ns
64 s subints
15 sub-bands
5 ns Jitter limit in
an hour
Best nanohertz
gravitational
wave radiometer
we know of
(Funky orbital
variations)



The double pulsar at new and “gently used” telescopes



Credit: D. Reardon



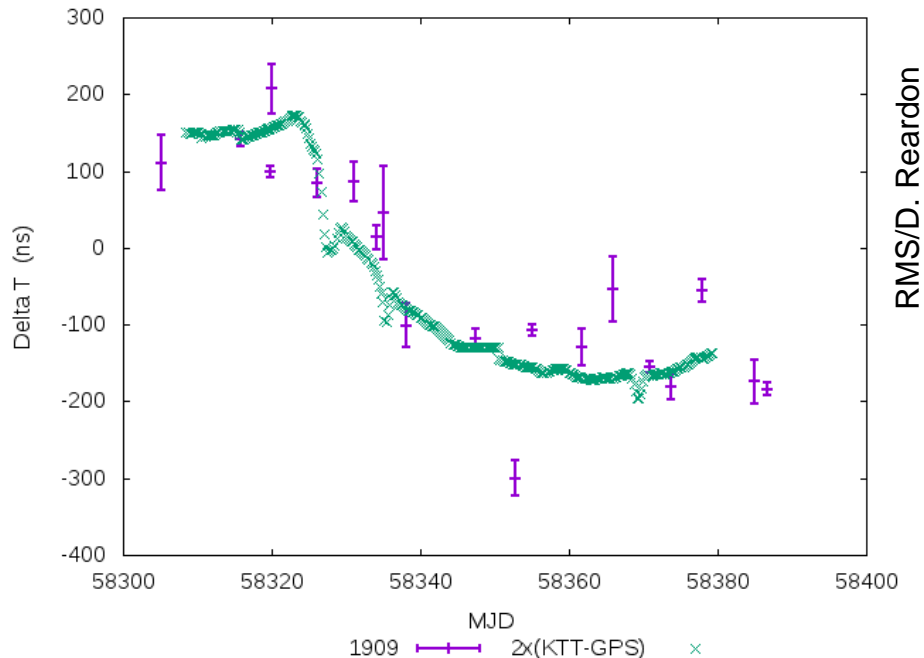
UTMOST+PKS public data: Credit: M. Lower



MeerKAT

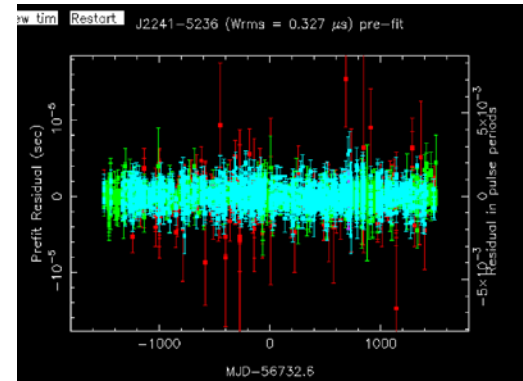
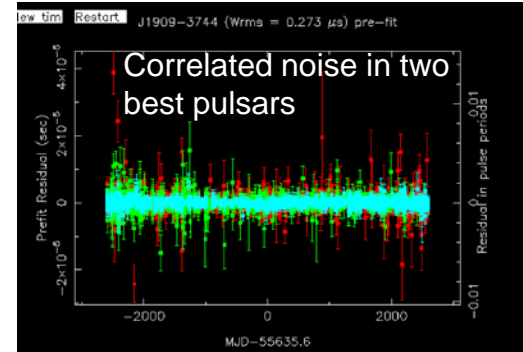
- Impressive raw sensitivity
- World-beating clock system
- Commissioning works in progress (Polarization calibration/channelization and firmware)
- Telescope access (will we get 5000 hrs)
- How does MeerKAT fit into IPTA efforts?

Commissioning data on J1909-3744 can see the very small MeerKAT clock drift



Parkes Pulsar Timing Array: The Future

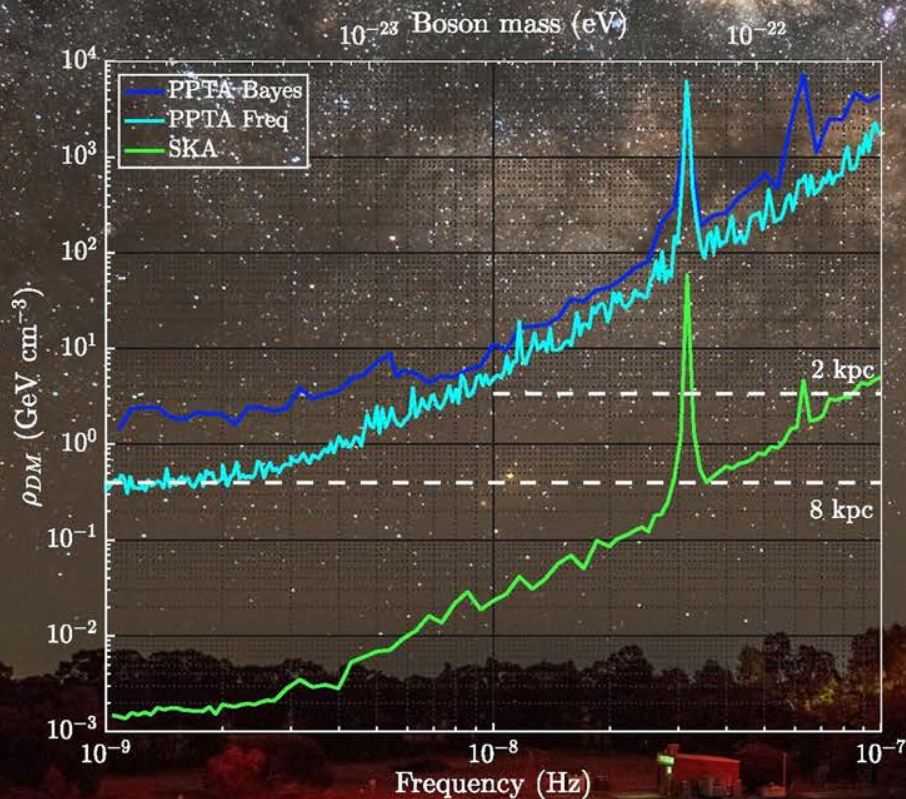
- Ultra-wideband system: 0.7 - 4 GHz
 - Need new methods to time-tag pulses
 - Common machinery will be needed for MeerKAT
- Data sets:
 - Approaching a second data release
 - Curation essentially complete
 - Timing analysis ongoing (in OzStar queue...)
 - GW search by Sunday?
- Vision:
 - How does PPTA fit in w.r.t IPTA and meerkat?



Credit: D. Reardon



PPTA constraints on fuzzy dark matter



✓ Physical Review D Editor's Suggestion

- PPTA DR19, >20 pulsars for > decade
- Constrain dark matter density in the Earth vicinity; $< 6 \text{ GeV cm}^{-3}$ for $m < 10^{-23} \text{ eV}$ --> probably not all dark matter is made from ultralight bosons
- Improve on previous results by a factor of ~ 5
- High cadence timing of Galactic Centre pulsars will produce the most promising results at the cosmologically favourable mass range $m \sim 10^{-22} \text{ eV}$

Young Pulsar Timing

Regular monitoring of **260 young and energetic pulsars** with the Parkes radio telescope having timing baselines of **~10 years**.

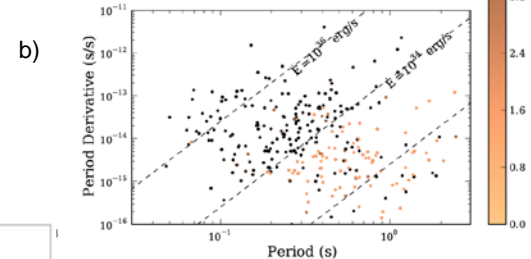
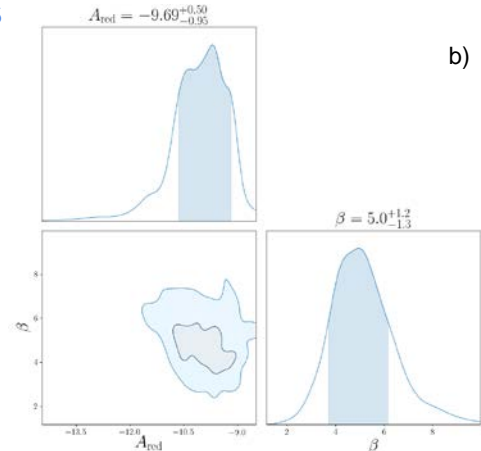
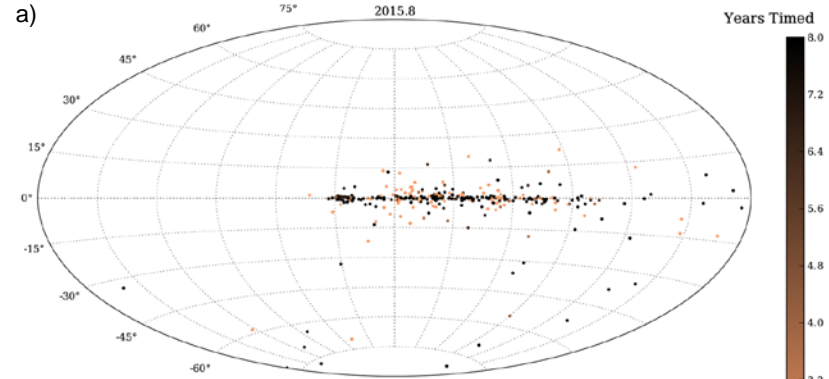
Provides several unique insights into neutron star (NS) science:

1. Rotational irregularities in NS: **Timing noise**
2. Probes into NS interiors: **Glitches**
3. Understanding NS spin-down: **Braking indices**
4. Astrometry: **Proper motion measurements**
5. Planetary companions: **Reflex motion**

The **Bayesian inference approach** to disentangle the various phenomena - applied to a large sample of young pulsars:

1. **New timing solutions.**
2. **Characterised timing noise in young pulsars.**
3. **5 new proper motions.**
4. **19 new braking index measurements.**

Future: The 1000 pulsar array with MeerKAT



a) Sample of young pulsars monitored by Parkes shown on the Galactic plane.
b) Distribution of Edot values for the same pulsar sample.

Credit: A.Parthasarathy

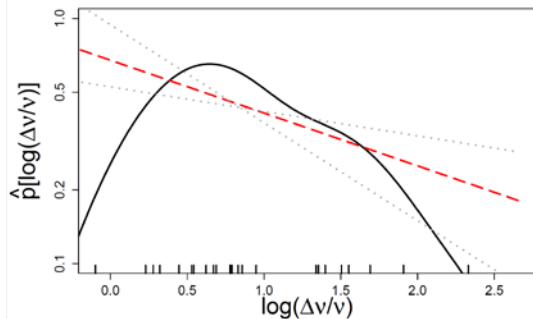
Integrated posterior distributions of red-noise amplitude and spectral index for a large sample of young energetic pulsars.



Glitches

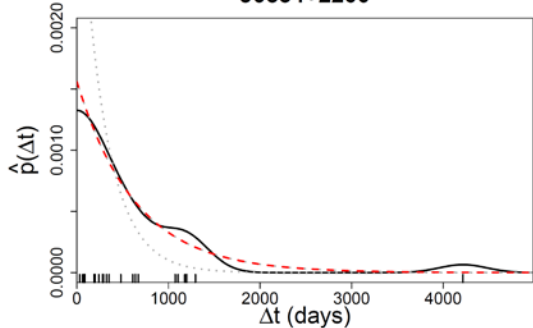
Amplitude

J0534+2200

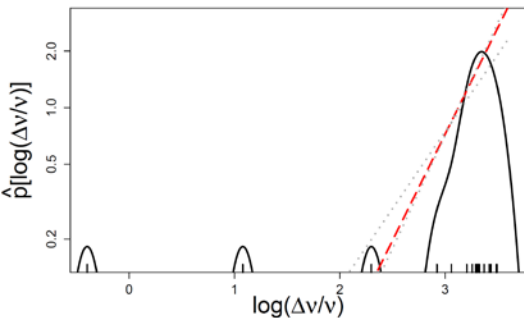


Waiting time

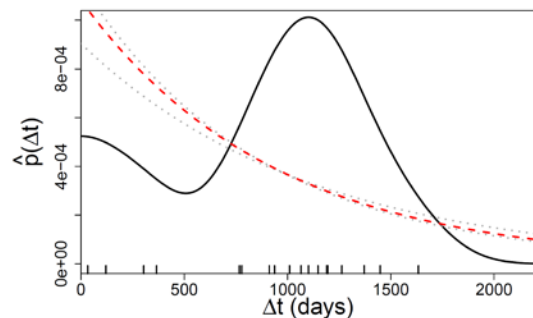
J0534+2200



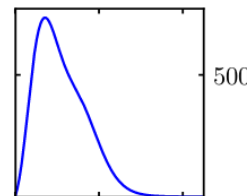
J0835-4510



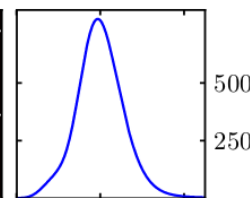
J0835-4510



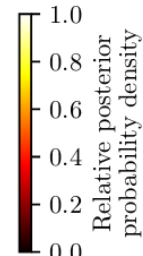
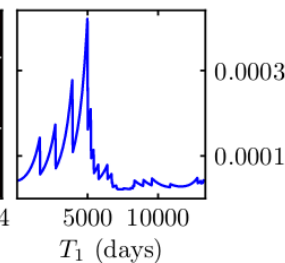
$p(\lambda_1)$



$p(\lambda_2)$



$p(T_1)$



λ_2 (days⁻¹)

λ_1 (days⁻¹)

λ_2 (days⁻¹)

T_1 (days)

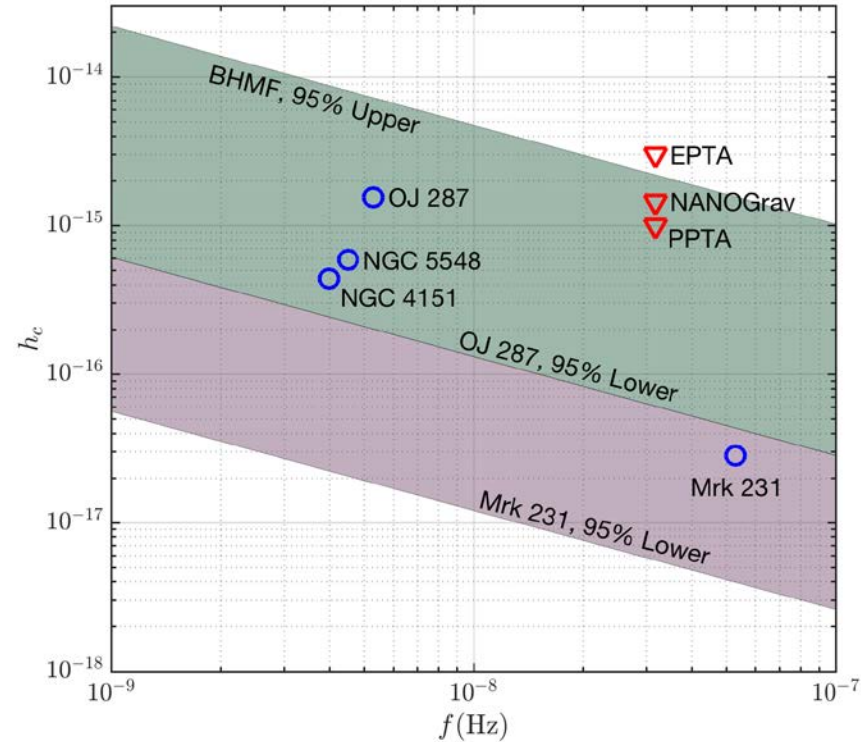
Size and waiting time statistics Howitt, Melatos & Delaigle 2018

Glitch correlations in the Crab
Carlin, Melatos & Vukcevic 2018



The minimum and maximum GWB from SMBBHs

- Zhu, Cui & Thrane MNRAS 2019
- The **maximum** signal $A_{\text{yr}} < 2.4 \times 10^{-15}$ with 95% C.L. with black hole mass function
- The **minimum** signal based on OJ 287: $A_{\text{yr}} > 6 \times 10^{-17}$ with 95% C.L.
- **3C 66B** ruled out as a SMBBH!
- $A_{\text{yr}} > 6 \times 10^{-18}$ with 95% C.L. if *at least one of OJ287, NGC 5548, NGC 4151 and Mrk 231 host a true SMBBH*
- A novel framework to quickly evaluate GWB implications for new SMBBH candidates/discoveries



Credit: X. Zhu



Plans for the retreat

- Start gravitational wave searching in PPTA dataset
- Collaboration: IPTA, PPTA, OzGrav, and MeerTime: how do they all fit together (with Chiara M.)
- Young pulsars stuff: connecting theory to observations
- Inference plans: Pulsar timing in Bilby, Profile-domain methods
- Pulsar searching

- Please contact us if you are interested in meeting this weekend.

- Key for success: identify OzGrav projects to contribute to international/domestic efforts
 - Swin, Monash, Melbourne Uni nodes



Plans for the next year

- Data sets
 - Complete PPTA data release
 - MeerKAT early results
 - J2241-5236
 - Young pulsar datasets: Glitching pulsars @ UTMOST/PKS
- Inference
 - Apply existing GW search codes (enterprise, PAL2, temponest, NX01?)
 - Developing GW search code (bilby)
 - Wide-band timing
 - Profile domain timing
 - Timing noise and glitch and modelling
- Pulsar Searching
 - Applying Viterbi continuous wave search to pulsar search

